



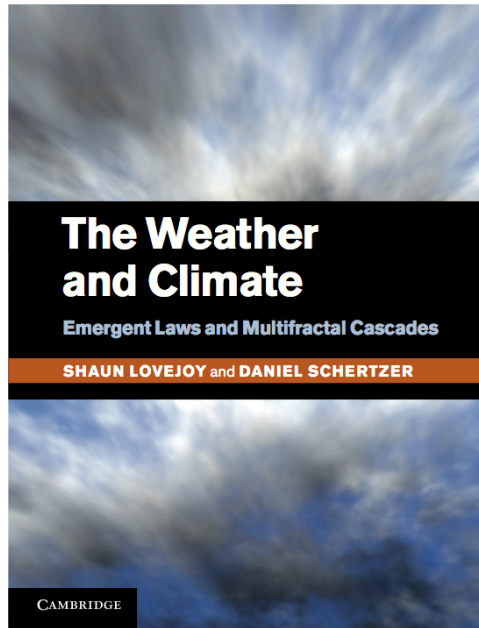
Scale, scaling and multifractals in complex geosystems part 1

Short course on: Scale, scaling and
multifractals in complex geosystems, EGU,
April 28, 2014, 17:30–19:00, Room B3

S. Lovejoy, U. McGill
D. Schertzer, U. de Paris Est

EGU Short Course 2014

Graduate course at McGill: Multifractals and turbulence



Course
synopsis

Today

12x2 hours, slides available:

[Lecture 1, Jan. 15, 2014](#), Introduction: Our multifractal world part 1

[Lecture 2, Jan. 22, 2014](#), Introduction: Our multifractal world part 2

[Lecture 3, Jan. 29, 2014](#), Turbulence and spectra

[Lecture 4, Feb. 5, 2014](#), Spectra, turbulence, fractal sets

[Lecture 5, Feb. 12, 2014](#), Fractal sets, multifractal cascades

[Lecture 6, Feb. 14, 2014](#), Multifractals: moments

[Lecture 7, Feb. 19, 2014](#), Data analysis

~~[Lecture 8, March 12, 2014](#), Multifractals: codimensions~~

[Lecture 9, March 19, 2014](#), Multifractals: extremes

[Lecture 10: March 26, 2014](#): Multifractal simulations

[Lecture 11: April 4, 2014](#): Generalized Scale Invariance: linear

[Lecture 12: April 9, 2014](#): Generalized Scale Invariance: nonlinear,
space-time

[http://
www.physics.mcgill.ca/
~gang/PHYS616/
PHYS616.home.htm](http://www.physics.mcgill.ca/~gang/PHYS616/PHYS616.home.htm)

EGU Short Course 2014



The unity of clouds
and rocks:

Scaling

Multifractal simulation

Emergent scaling laws

$$\text{Fluctuations} \approx (\text{turbulent flux}) \times (\text{scale})^H$$

Differences,
tendencies,
wavelet
coefficients

Cascading
driving flux

Anisotropic
Space-time
Scale function

Fluctuation
/conservation
exponent

Fourier domain:

$$\left(\frac{\text{Variance}_{\text{observables}}}{\text{wavenumber}} \right) = \left(\frac{\text{Variance}_{\text{flux}}}{\text{wavenumber}} \right) (\text{wavenumber})^{-2H}$$

$$= (\text{wavenumber})^{-\beta}$$

Space: $E(k) \approx k^{-\beta}$

Time: $E(\omega) \approx \omega^{-\beta}$

Energy Spectra

Scaling geometric sets of points = fractals

Scaling fields = multifractals

$$E(k) \propto k^{-\beta}$$

$E(k)$ = "scaling" ← β = scale invariant"

$k = 2\pi/L$ = wavenumber, β = spectral exponent

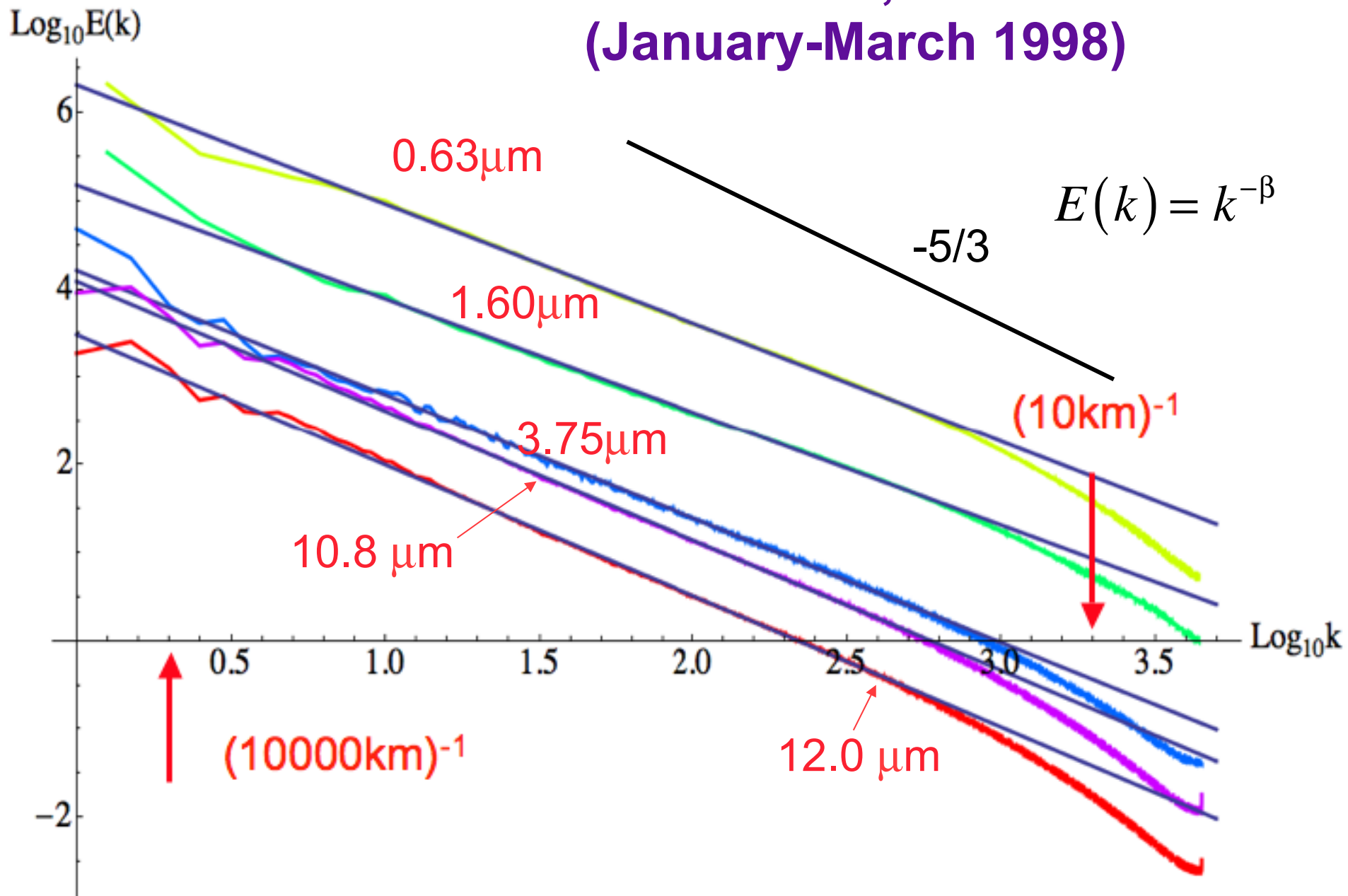
Scale invariance

$$E(\lambda^{-1}k) = \lambda^{\beta} E(k)$$

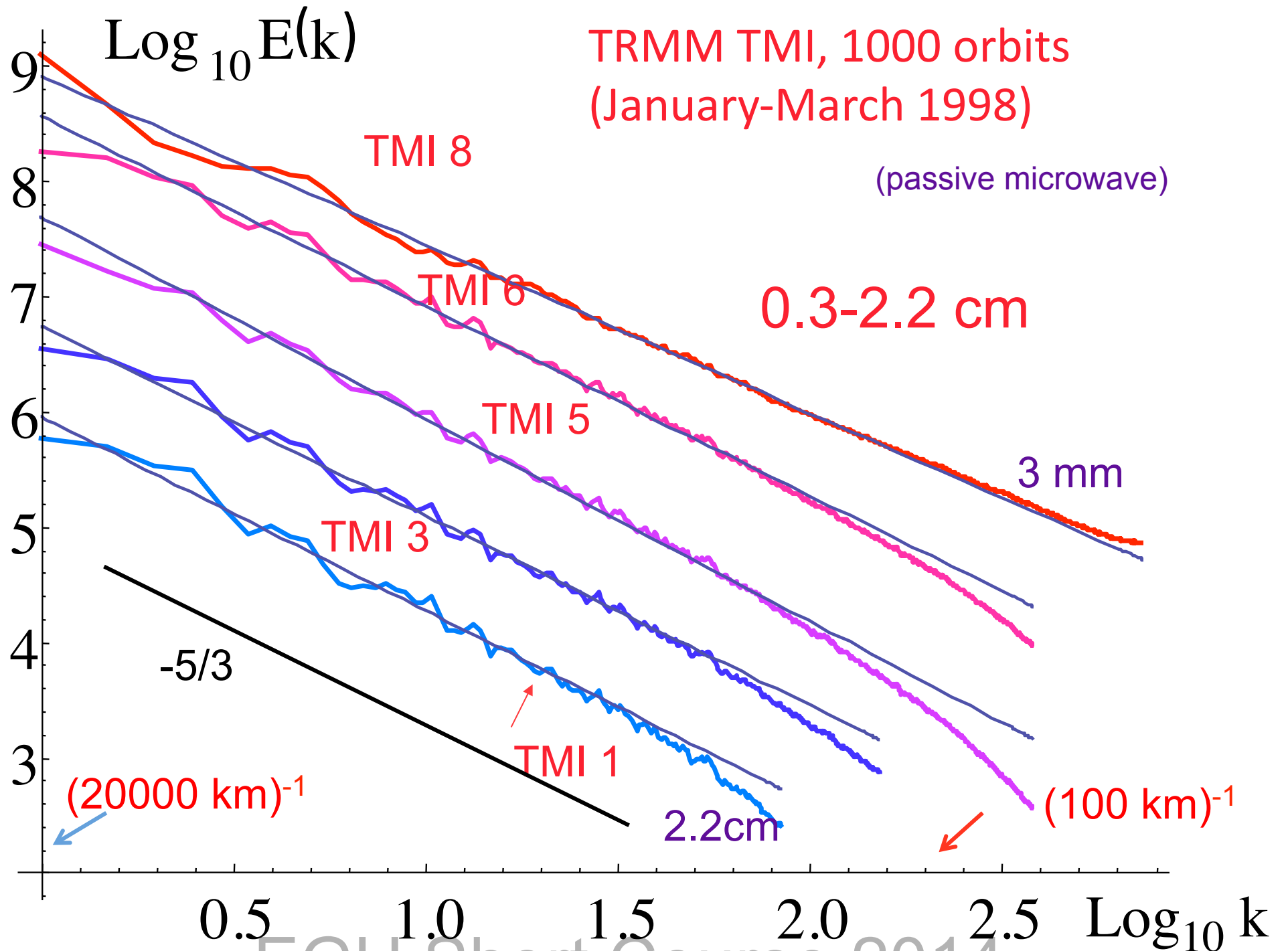
β Invariant under zoom by factor λ in real space.

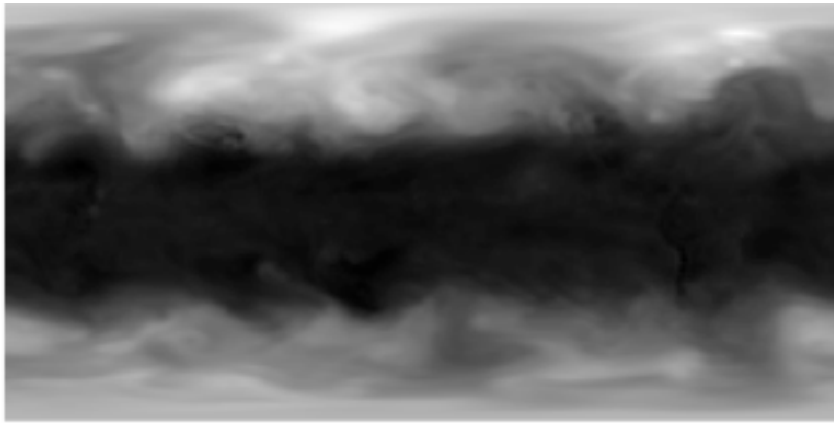
Scaling in space, a guided tour

TRMM VIRS, 1000 orbits (January-March 1998)



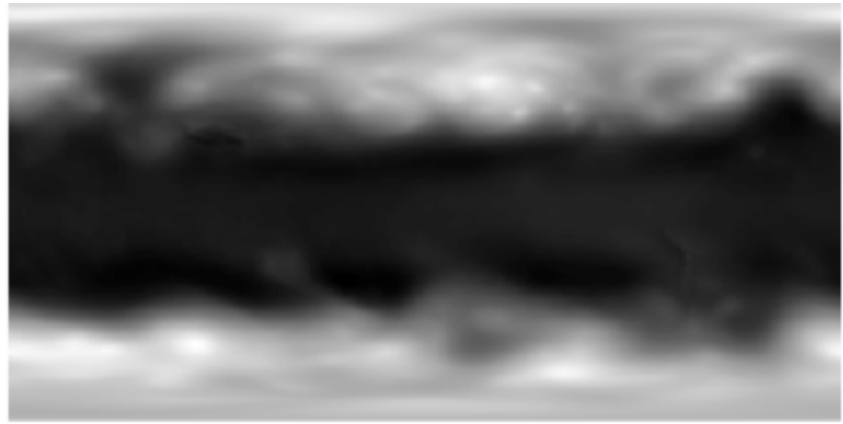
Visible, near infra red, thermal infra red



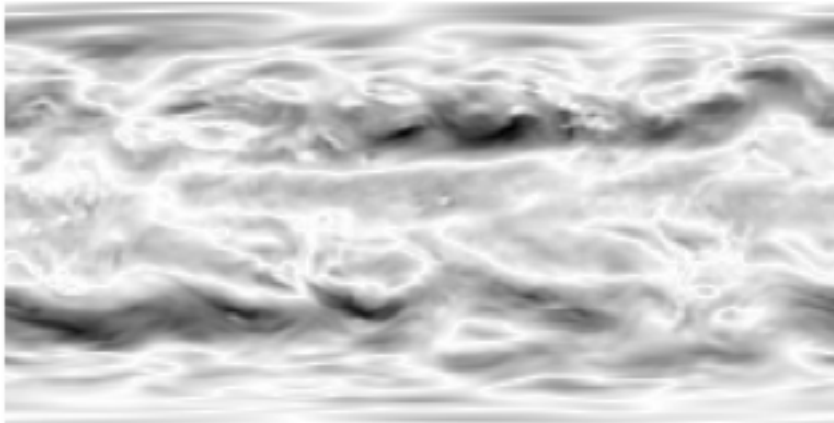


h

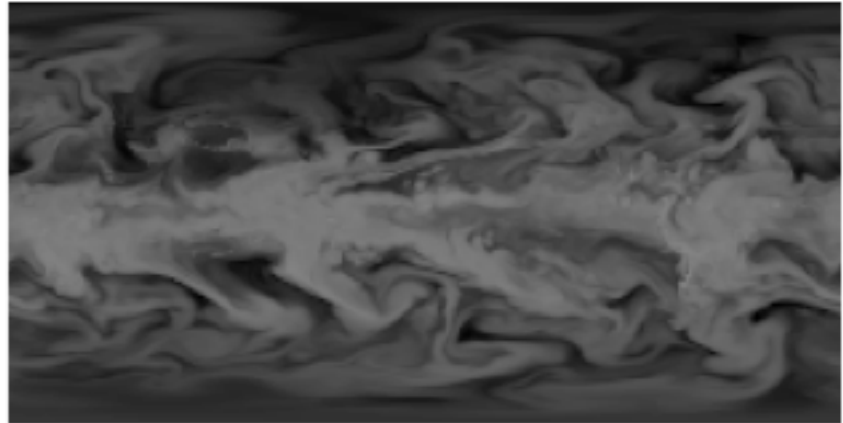
1.5a:



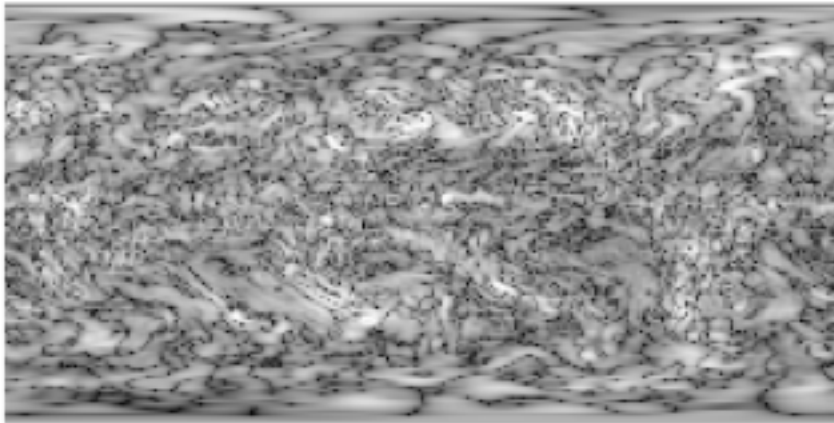
T



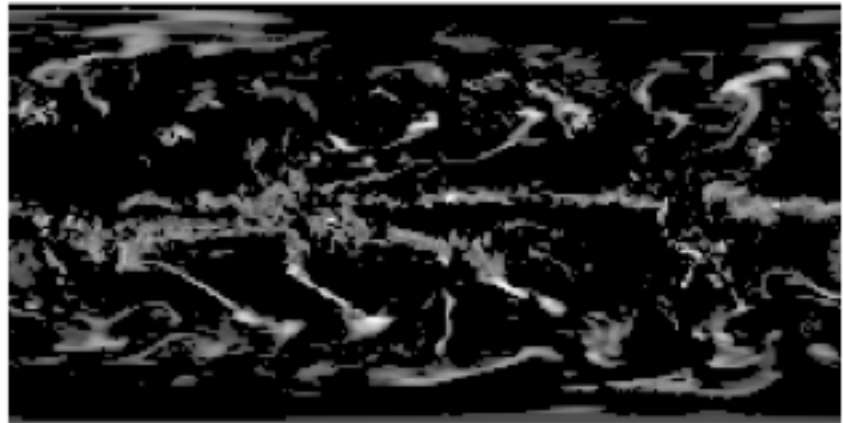
u



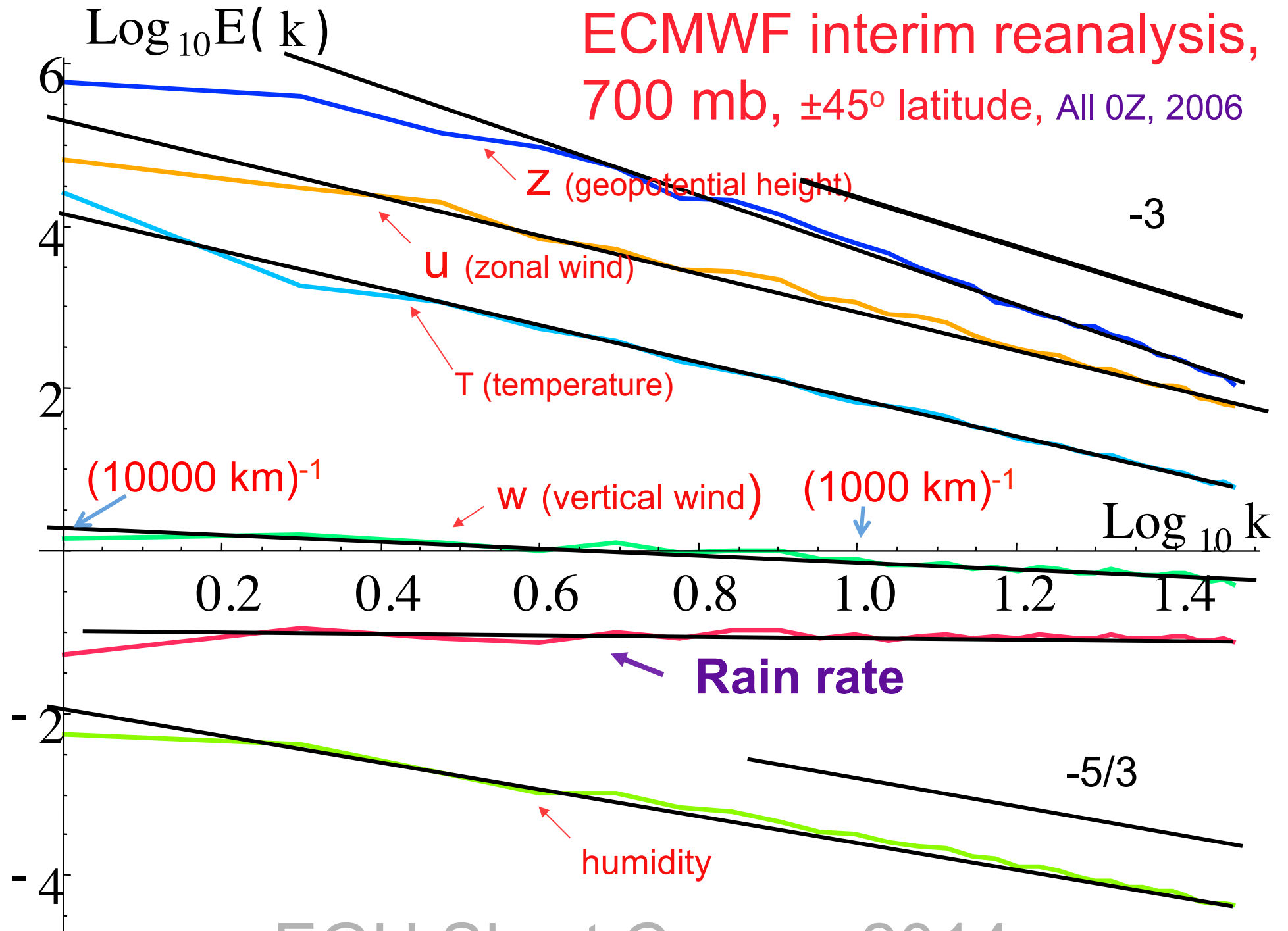
v



w

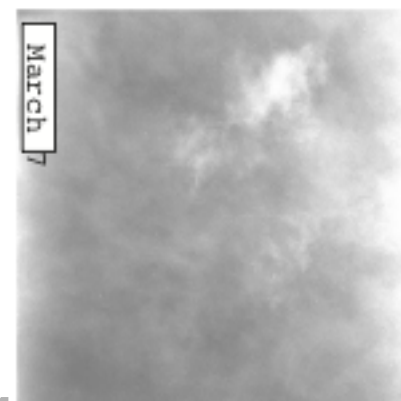
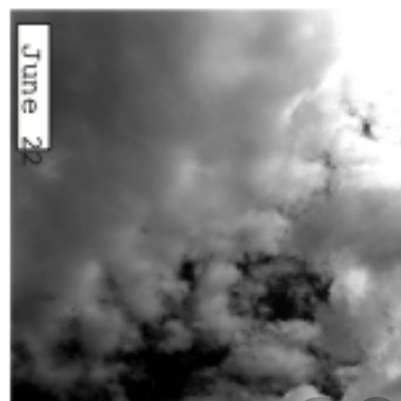
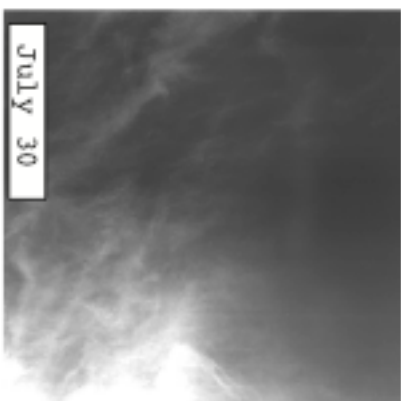
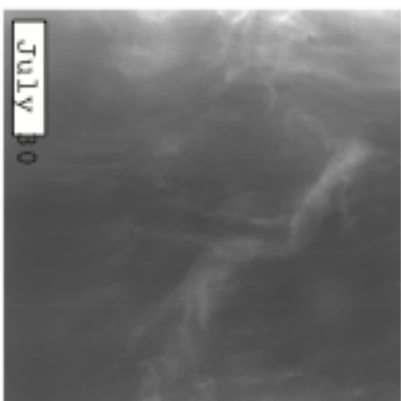


z

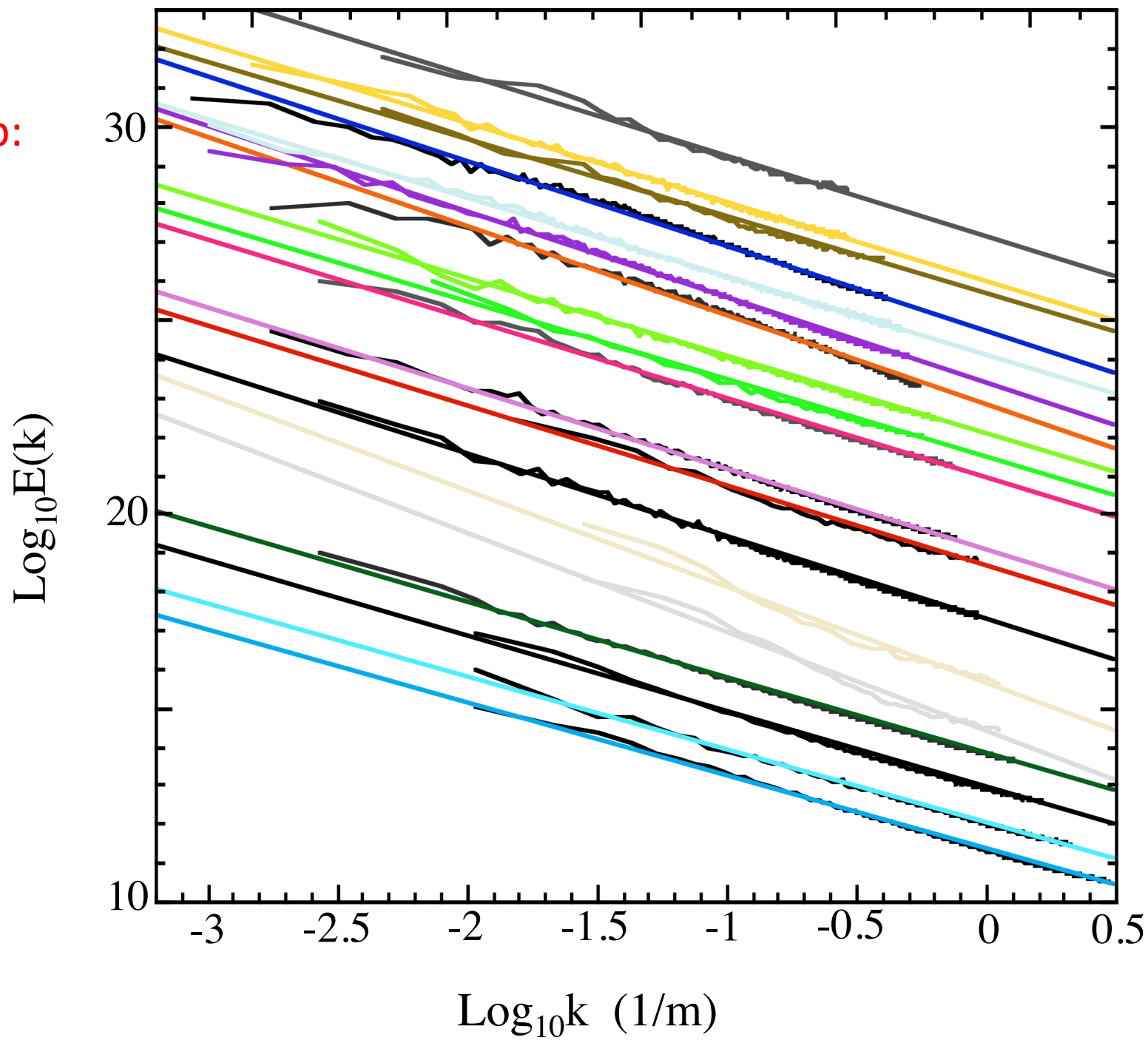




1.4a:



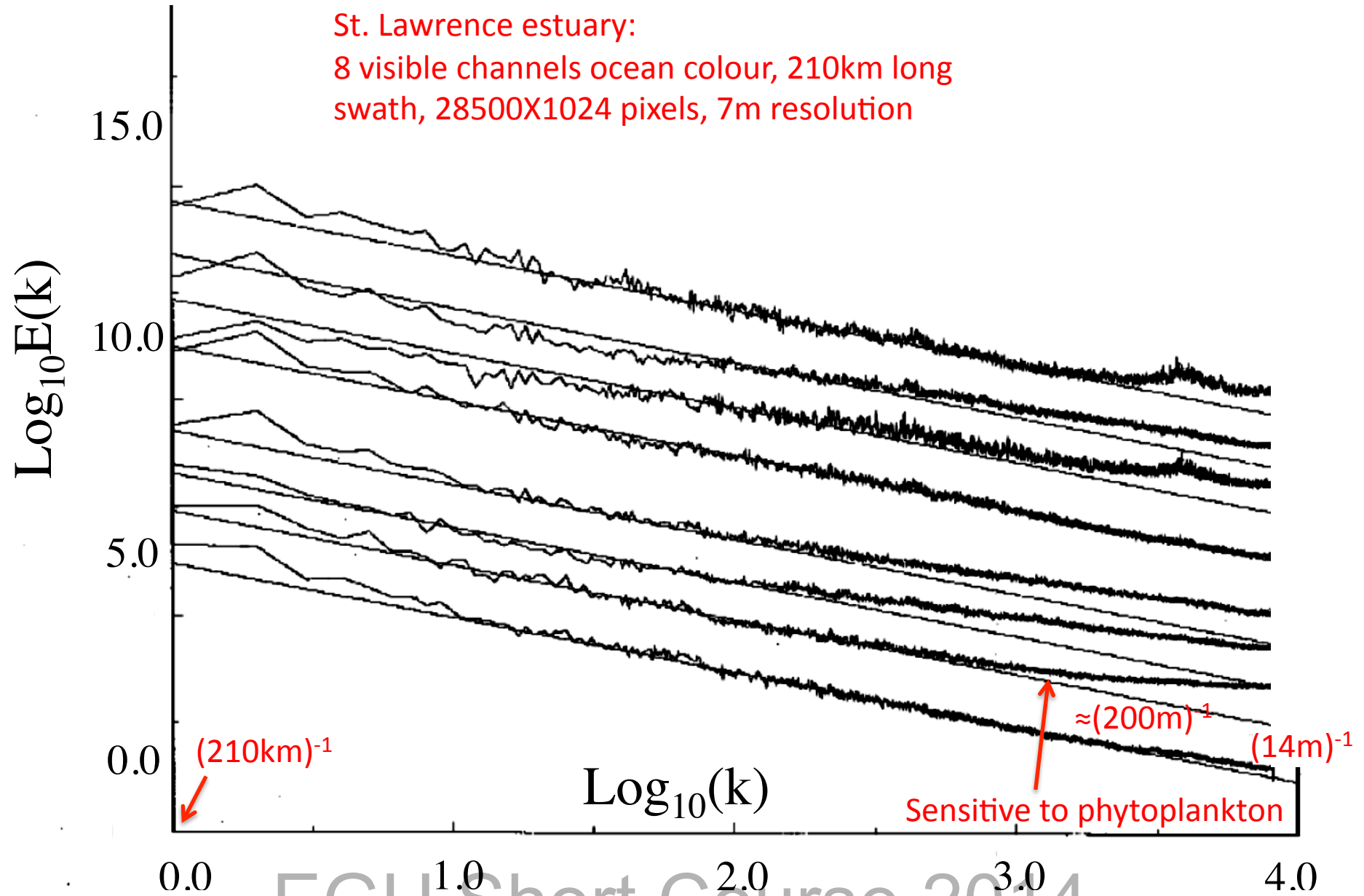
1.4b:



Ocean surface

St. Lawrence estuary:

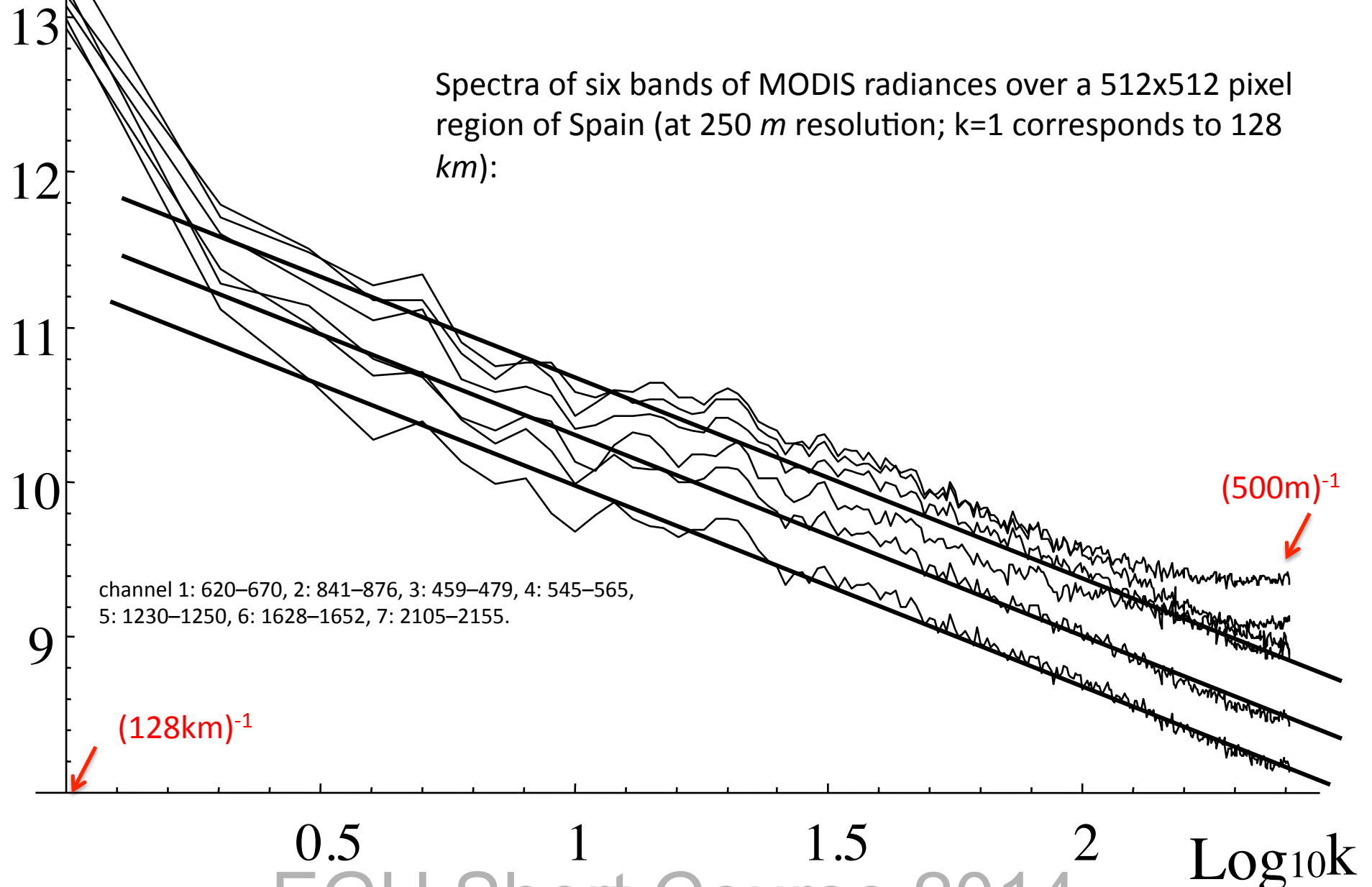
8 visible channels ocean colour, 210km long
swath, 28500X1024 pixels, 7m resolution



$\text{Log}_{10} E(k)$

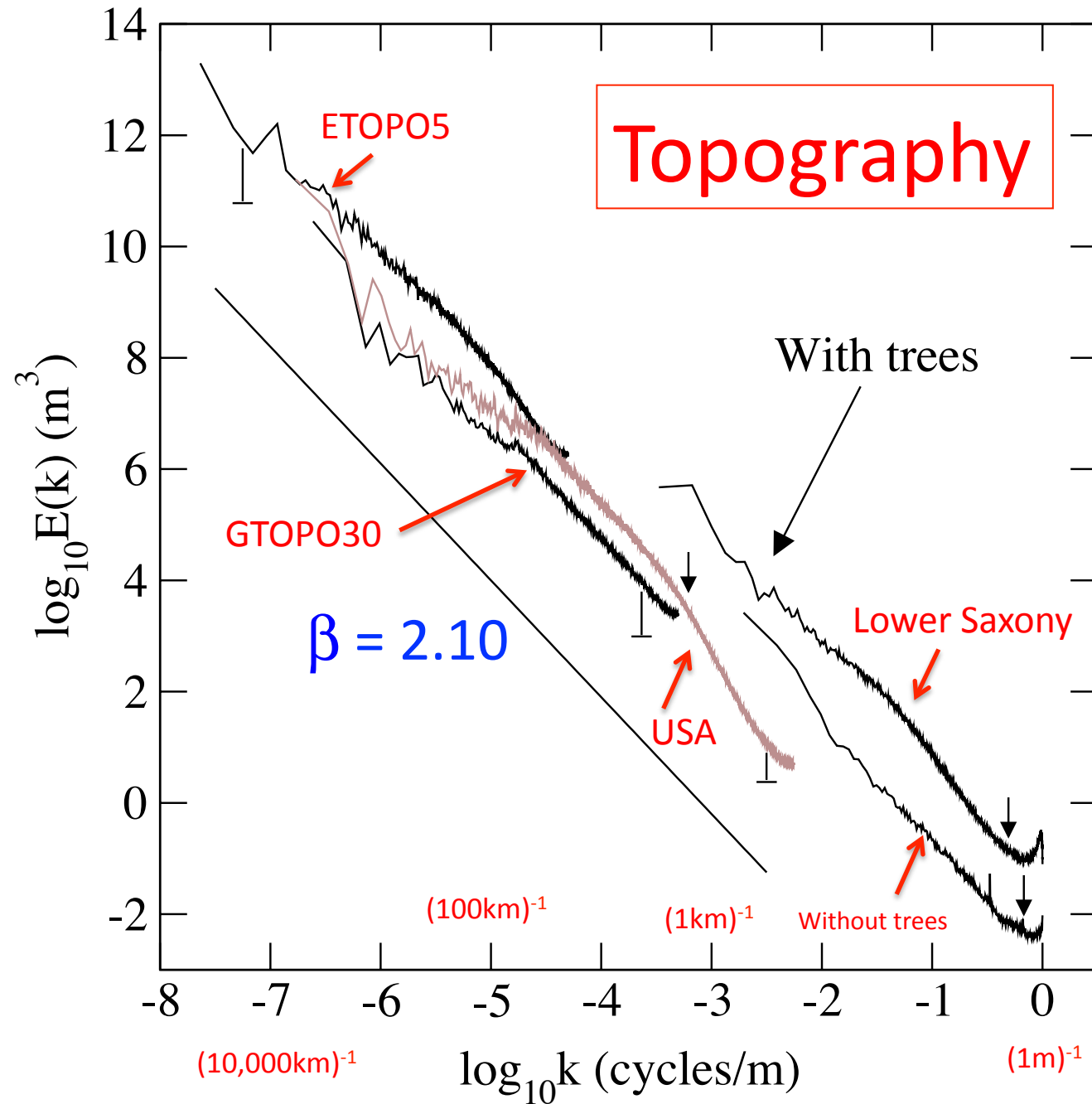
Vegetation and soil moisture indices

Spectra of six bands of MODIS radiances over a 512x512 pixel region of Spain (at 250 m resolution; k=1 corresponds to 128 km):





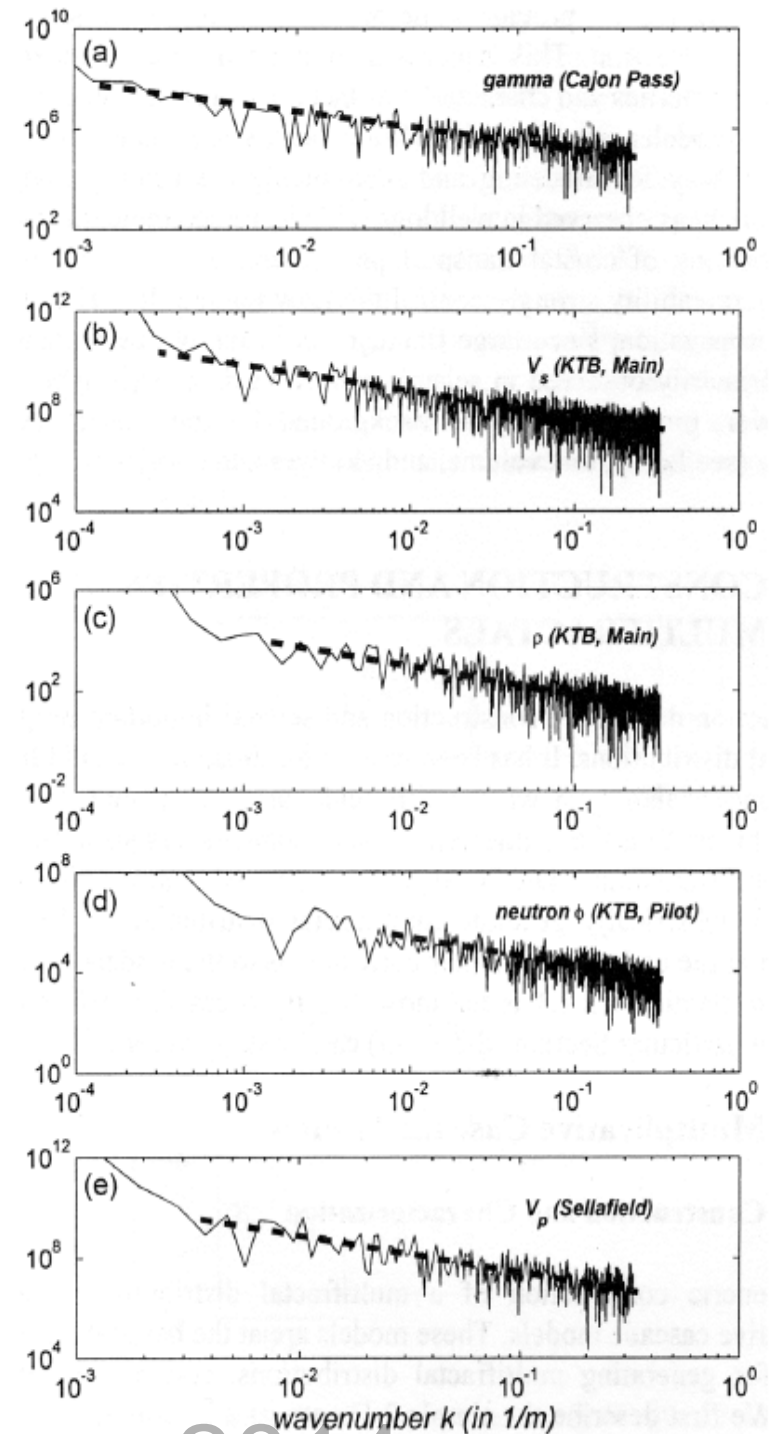
Topography



Gagnon, Lovejoy and Schertzer, 2006

The scaling of the KTB borehole

Marsan and Bean (2003)

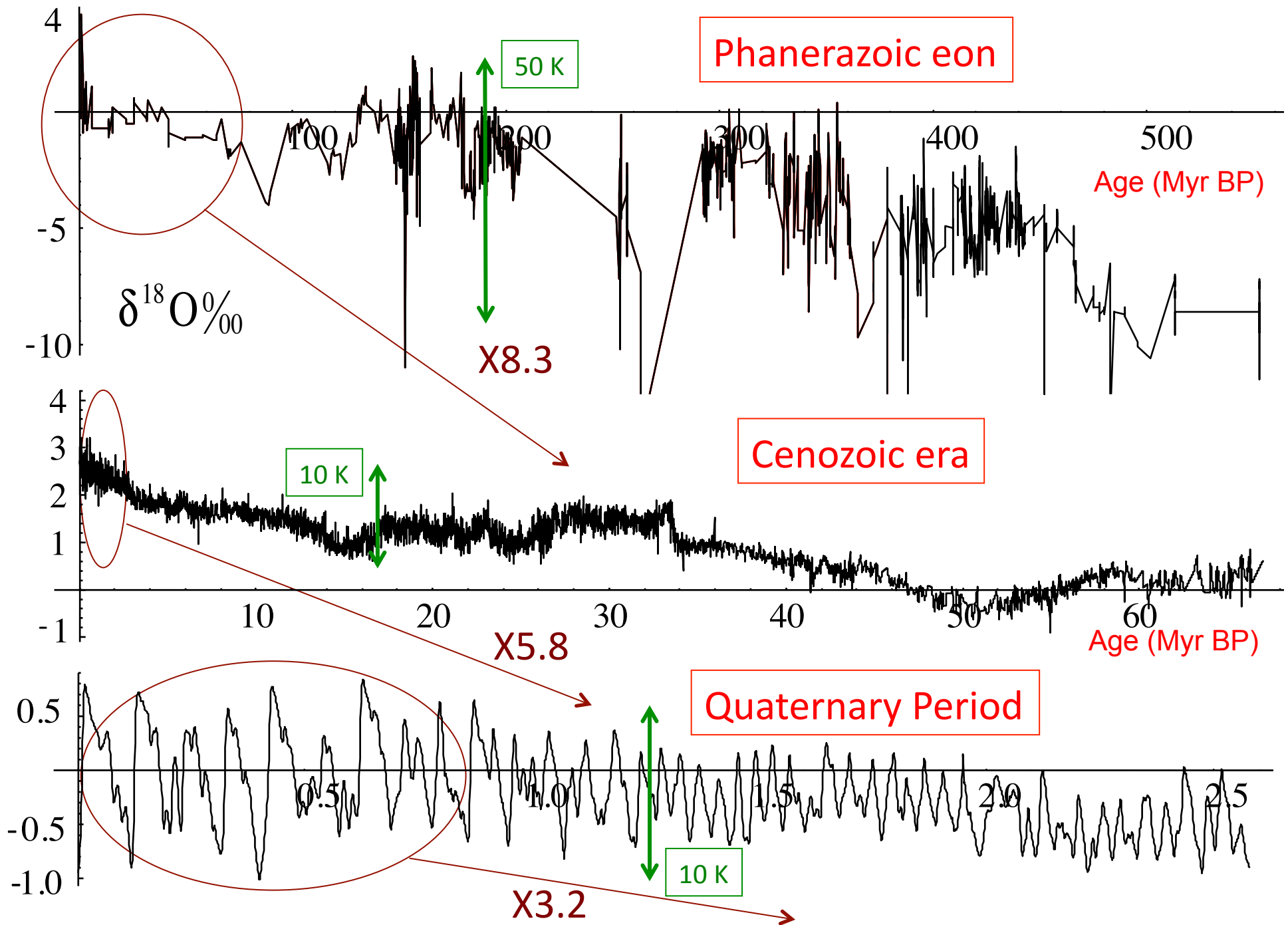


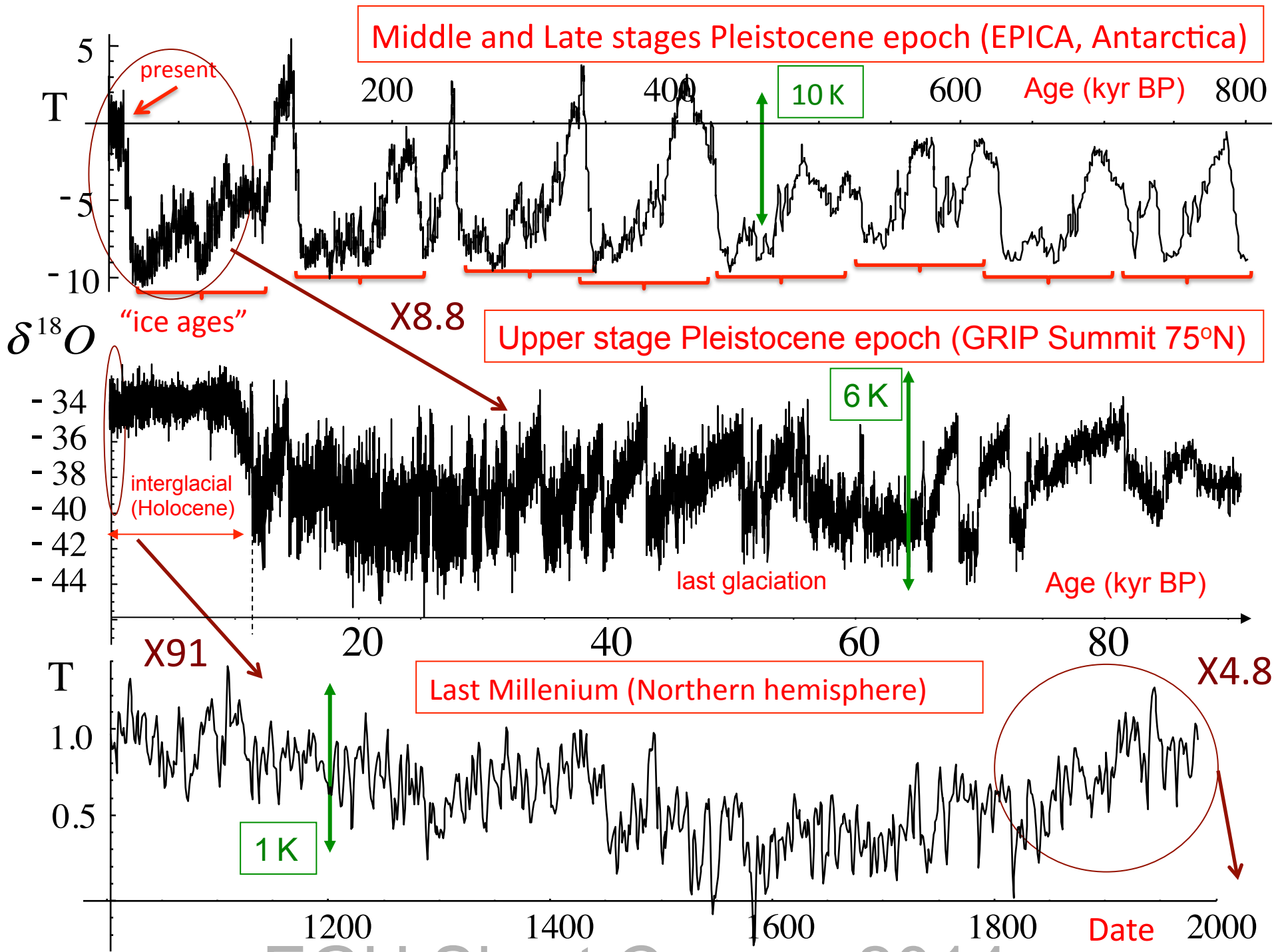
Scaling in time:

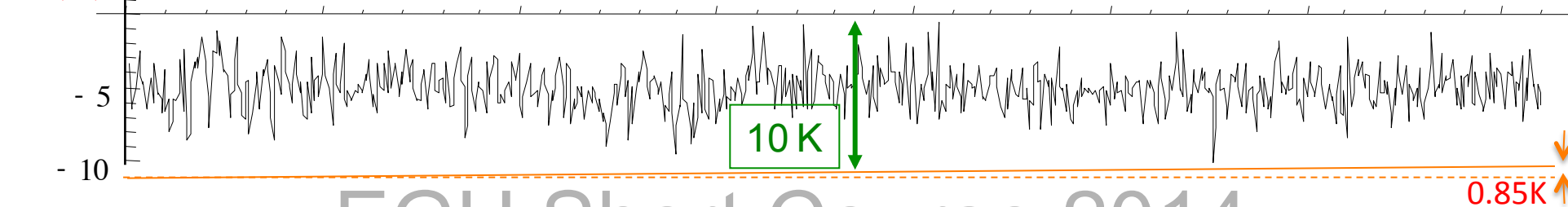
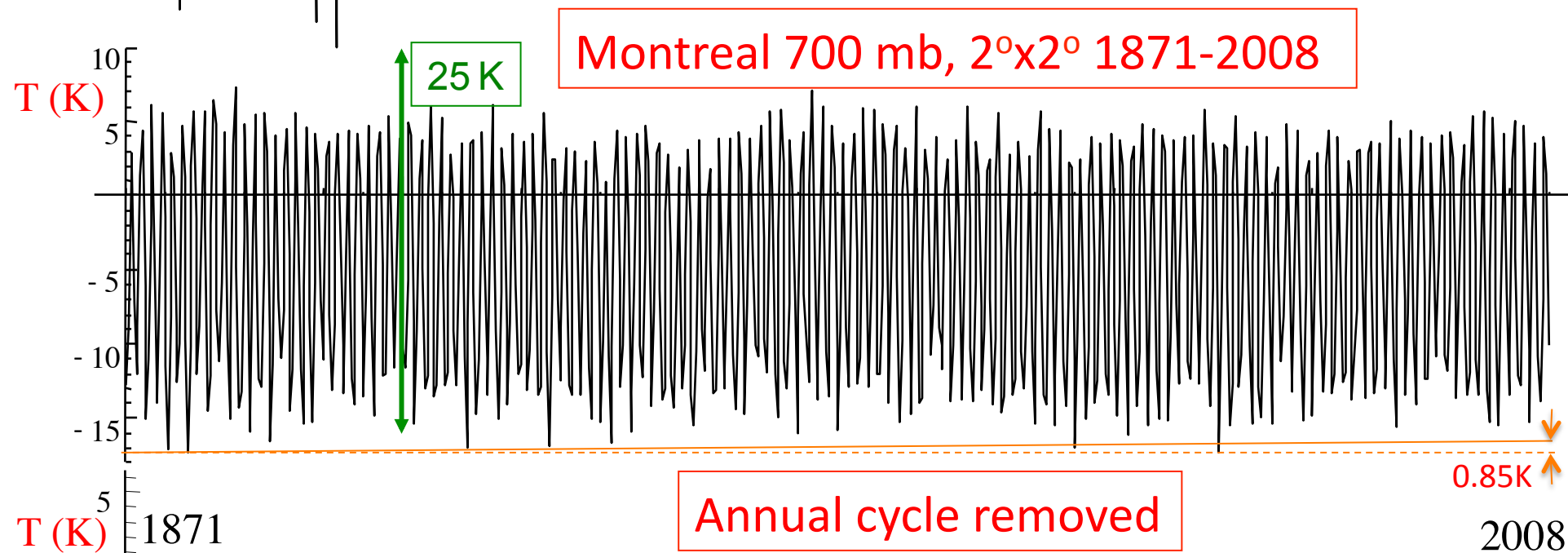
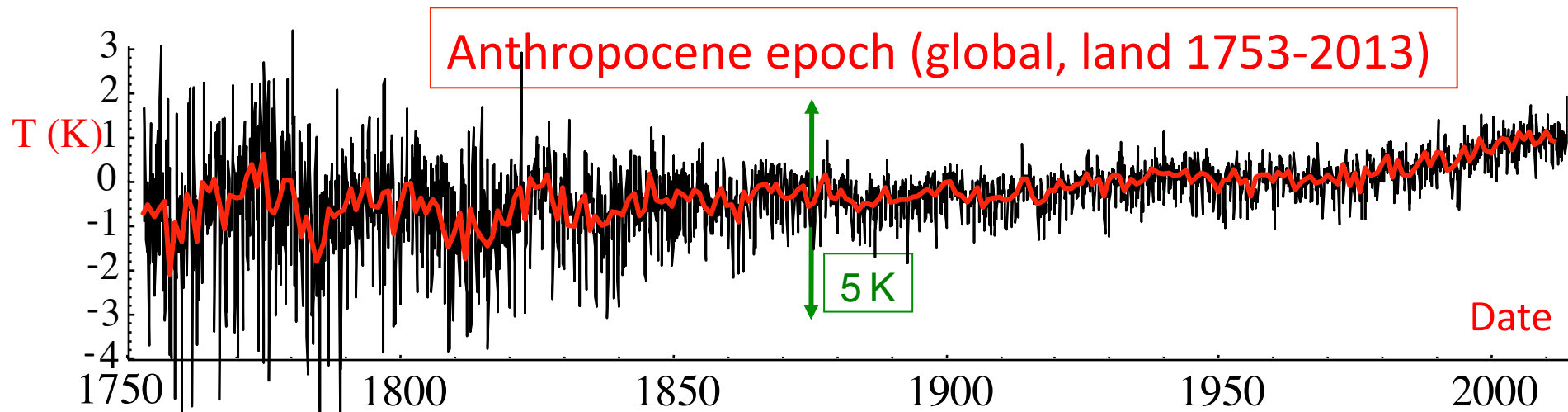
From the age of the earth to the
viscous dissipation scale: 4.5×10^9
years - 1 ms:

20 orders of magnitude in time

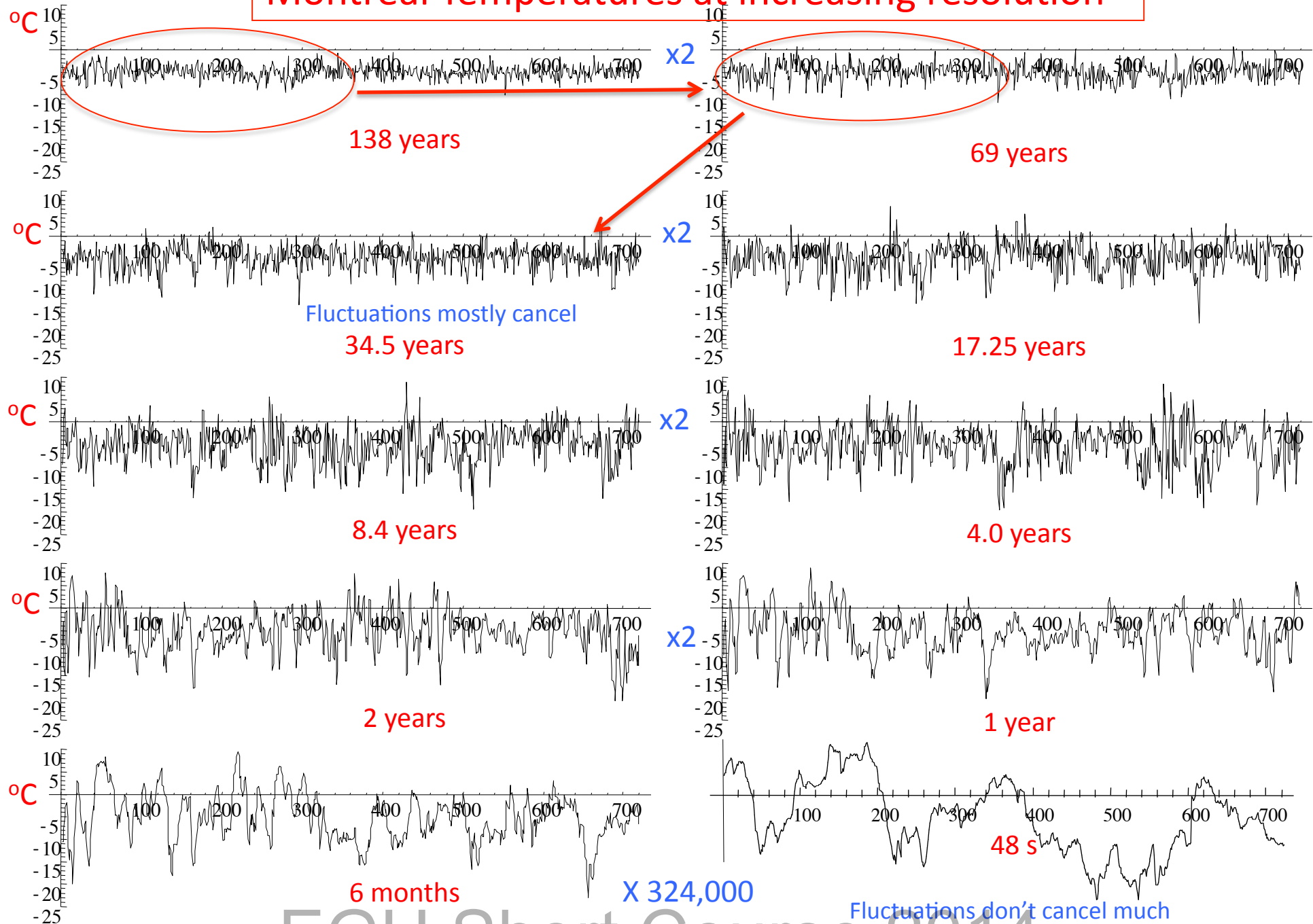
A voyage through scale...







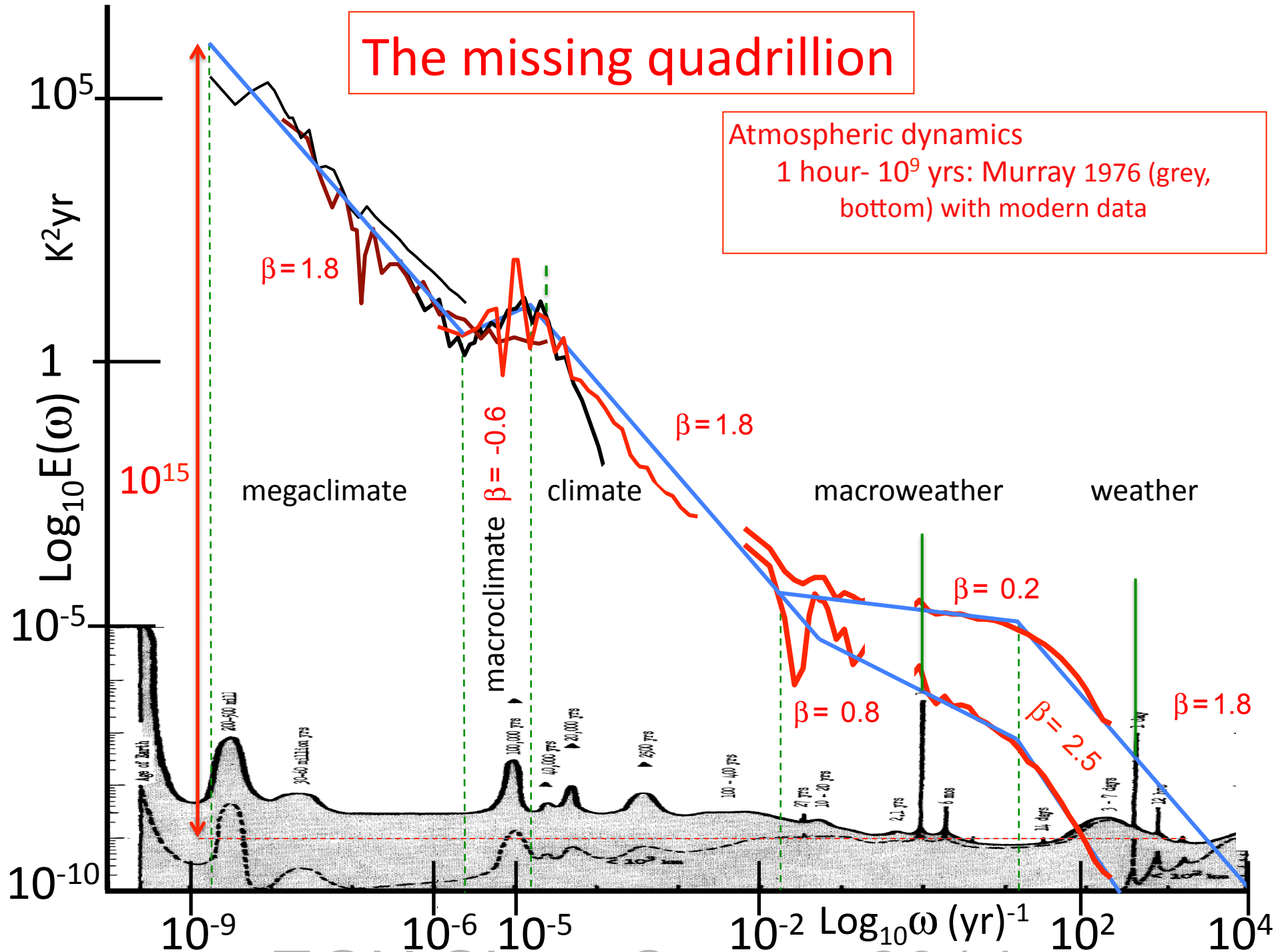
Montreal Temperatures at increasing resolution



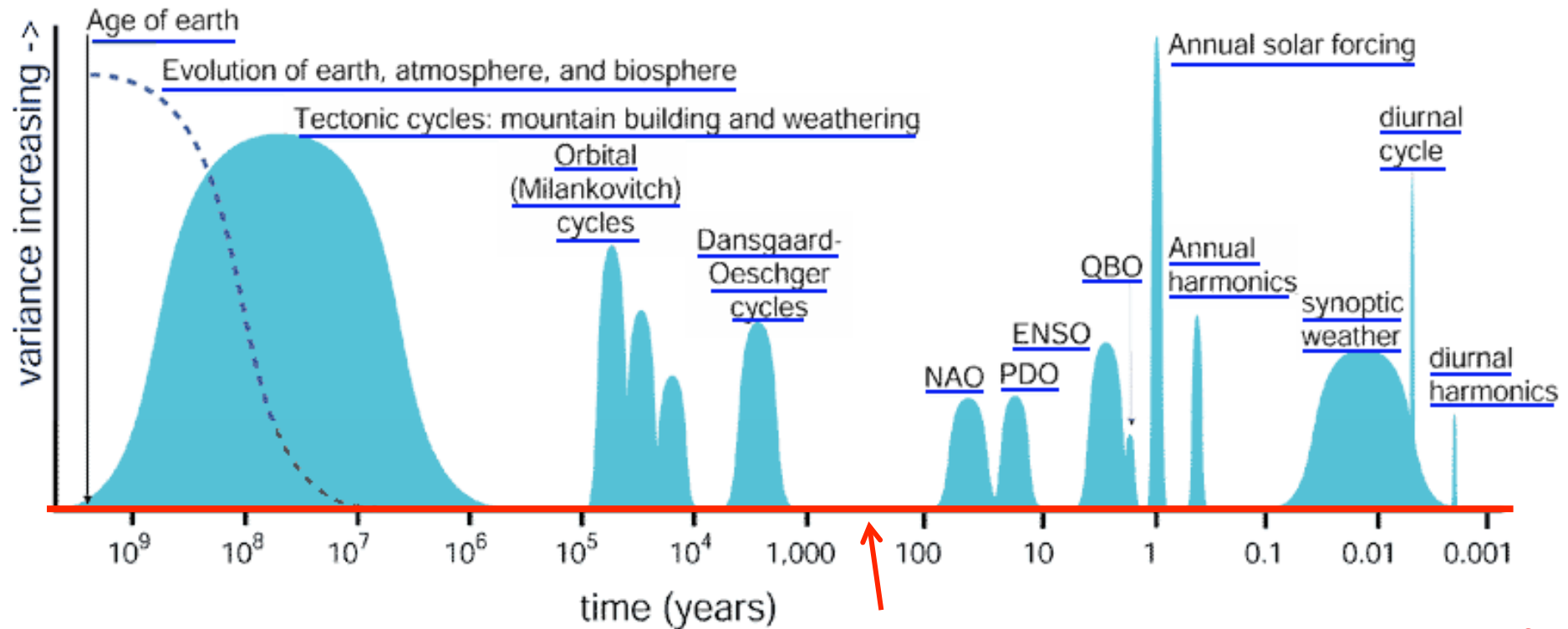
The missing quadrillion

Atmospheric dynamics

1 hour- 10⁹ yrs: Murray 1976 (grey, bottom) with modern data



The NOAA NCDC Paleoclimate data site graph (inspired by Mitchell)



The background is totally flat: error of $\approx 10^{16}$

The explanation of the figure:

"... figure is intended as a mental model to provide a general "powers of ten" overview of climate variability, and to convey the basic complexities of climate dynamics for a general science savvy audience."

The site assures us that just "because a particular phenomenon is called an oscillation, it does not necessarily mean there is a particular oscillator causing the pattern. Some prefer to refer to such processes as variability."

How to understand the variability?

Answer #1:

Scale bound thinking

Scale bound thinking

Antonie van Leeuwenhoek
(1632–1723)

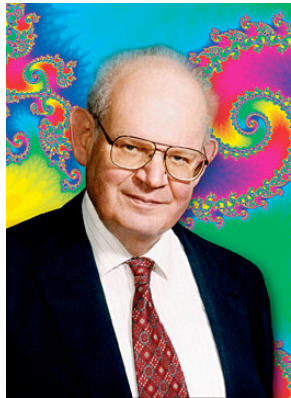


A new world in a drop of water

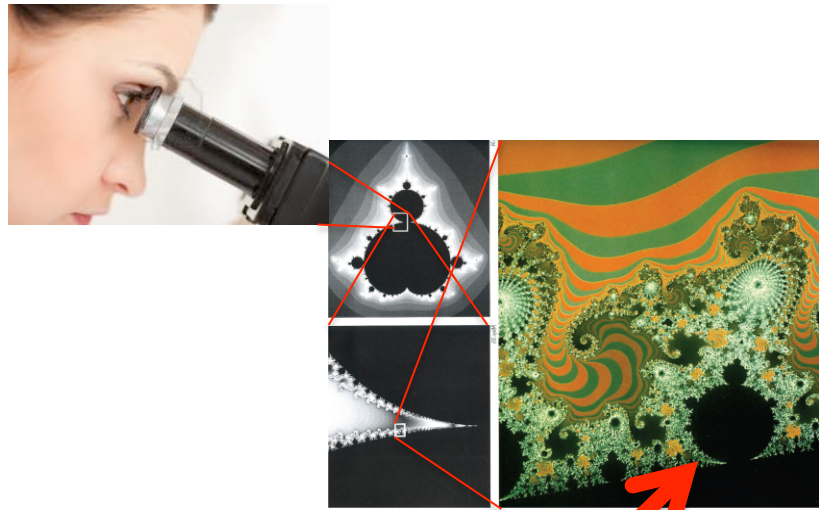
....the discovery of micro-organisms

"Animalcules," described in depth by Leeuwenhoek, c1695–1698. By Anton van Leeuwenhoek

Pure, (self-similar) Fractal thinking



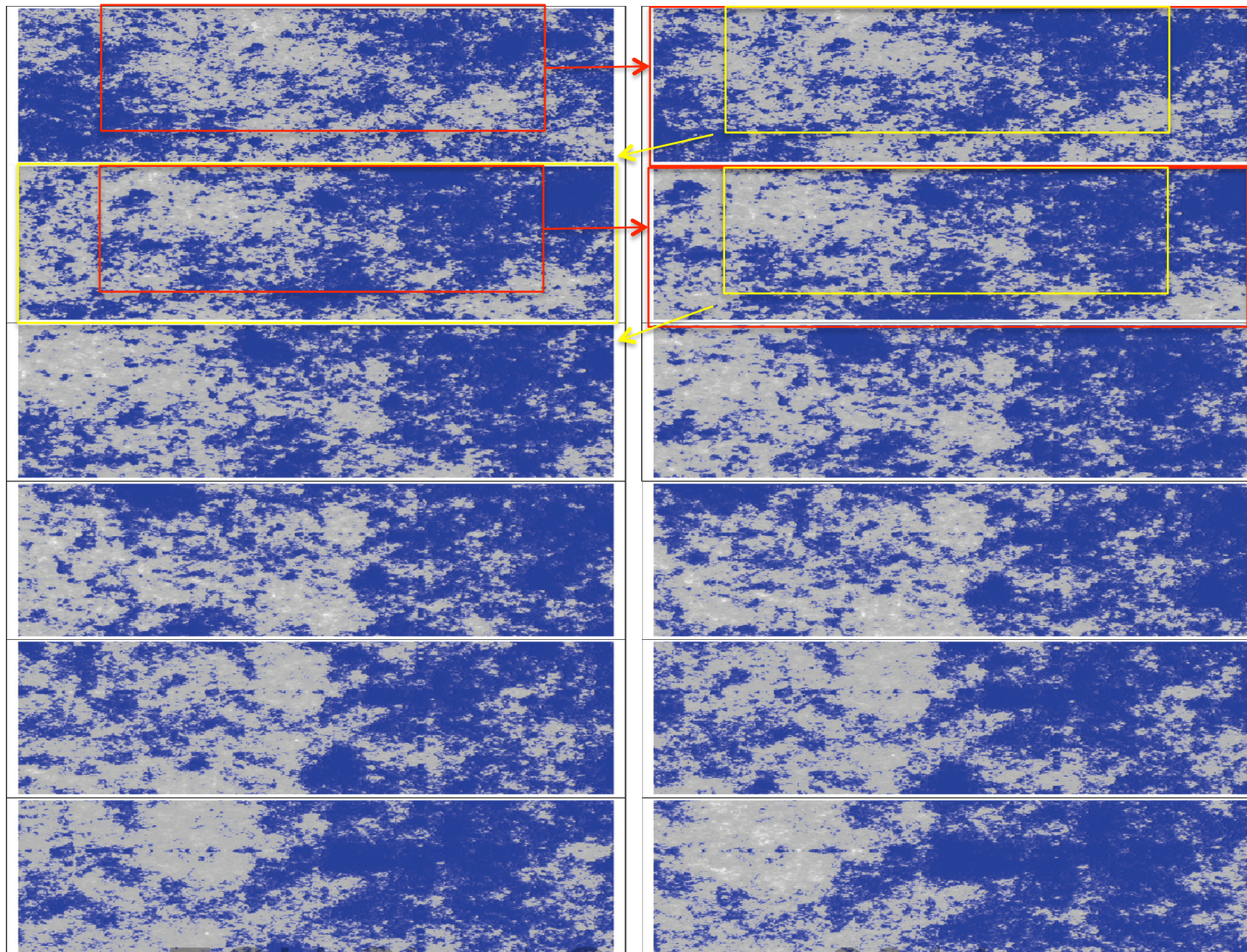
Mandelbrot 1924-2010



The same!!!

(the Mandelbrot set)

Self-similar Fractal thinking is OK here (Zooming in by factors of 1.7)





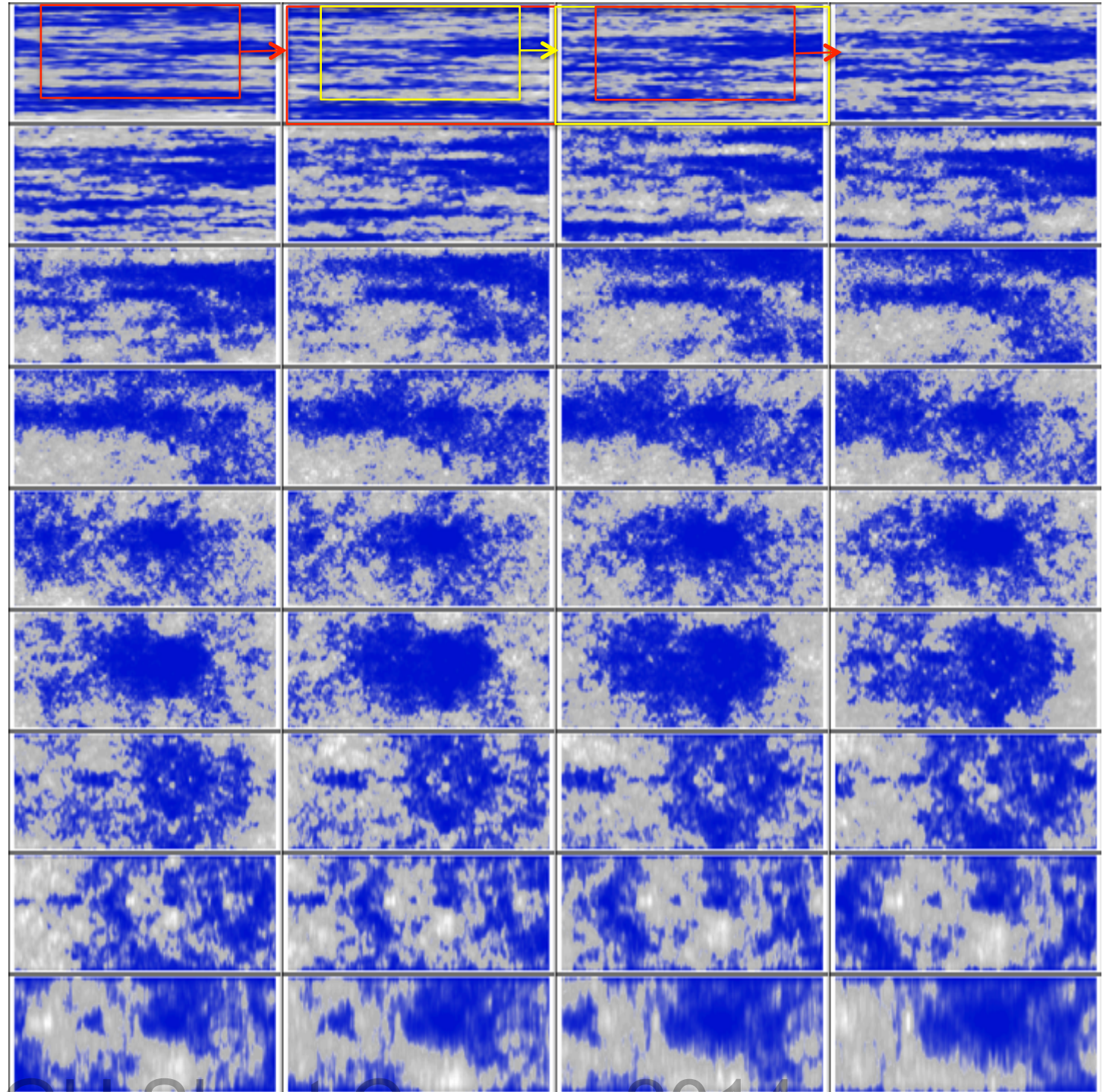
But not here!

**Need
Scale
invariant
thinking!**

(Zoom
factor 1000)



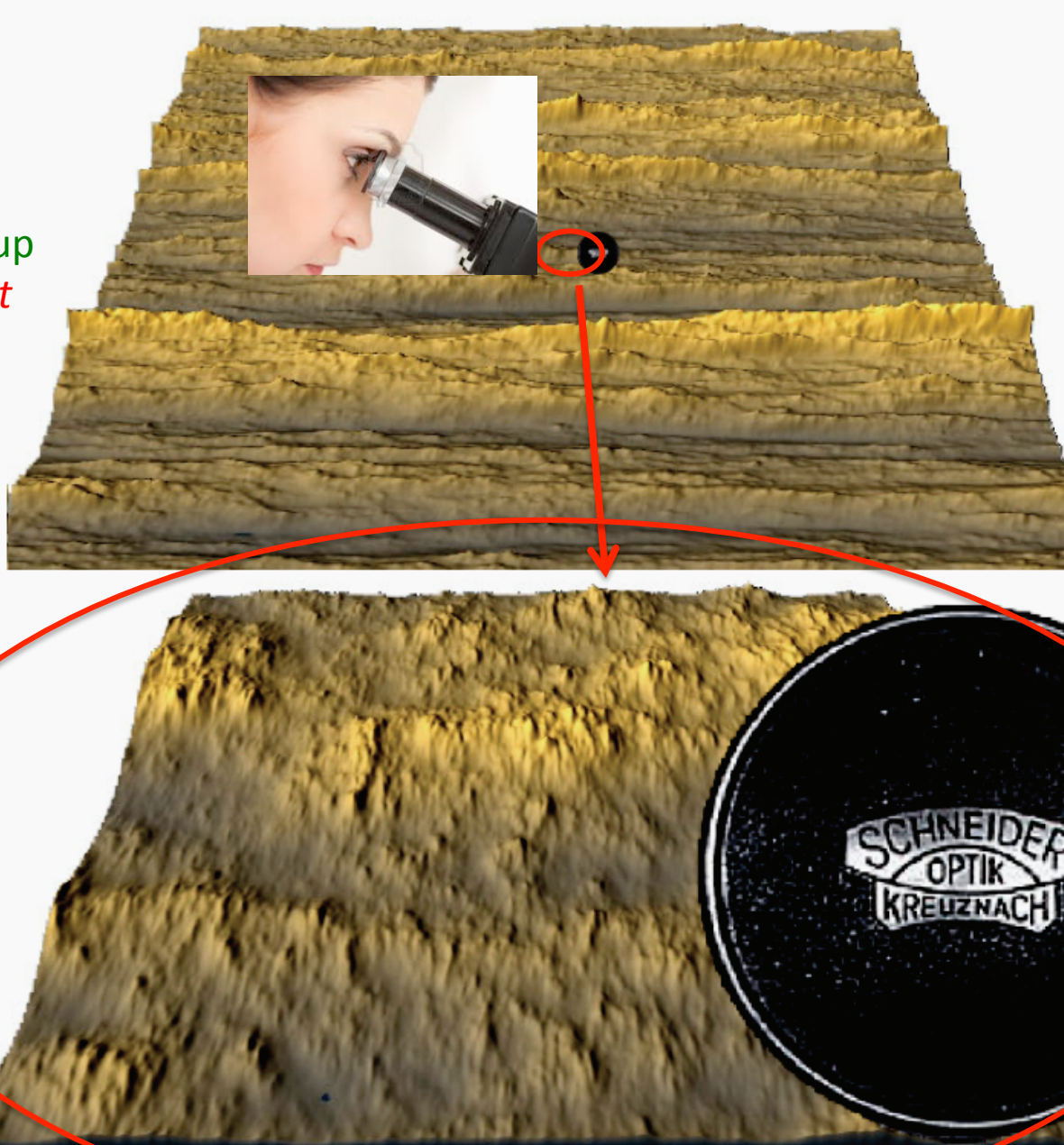
Vertical cross-section
of the atmosphere



Scale invariance and the Phenomenological Fallacy

- 1) Morphology not dynamics is taken as fundamental
- 2) Scaling is reduced to the isotropic (self-similar) special case

Isotropic Blow up
reveals *different*
morphology



Anisotropic multifractal surface simulation

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How to understand the variability?

Answer #2

- Scaling, scale invariance:

$$\Delta T(\Delta t) = \varphi \Delta t^H$$

Fluctuation

Time lag
(time scale)

Driving dynamic flux
(cascade, multifractal, see later)

Fluctuation exponent

Difference, tendency, Haar fluctuations

Differences: The difference in temperature between t and $t+\Delta t$

Tendency: The average of the temperature (with overall mean removed) between t and $t+\Delta t$

Haar: The difference between the average of the temperature from t and $t+\Delta t/2$ and from $t+\Delta t/2$ and $t+\Delta t$

Relations: When $1 > H > 0$: Haar \approx difference
When $0 > H > -1$: Haar \approx tendency

Fluctuations and wavelets

In wavelet analysis, one defines fluctuations with the help of a basic "" $\Psi(x)$ and performs the convolution:

$$\Delta v(\Delta t) = \frac{1}{\Delta t} \int v(t') \Psi\left(\frac{t-t'}{\Delta t}\right) dt'$$

mother wavelet

Difference

$$(\Delta v)_{diff} = v(t + \Delta t / 2) - v(t - \Delta t / 2)$$

$$\Psi(t) = \delta(t - 1/2) - \delta(t + 1/2)$$

Tendency

$$(\Delta v)_{tend} = \frac{1}{\Delta t} \int_t^{t+\Delta t} v'(t') dt'; \quad v'(t) = v(t) - \overline{v(t)}$$

$$\Psi(t) = I_{[-1/2, 1/2]}(t) - \frac{I_{[-\tau/2, \tau/2]}(t)}{\tau}; \quad \tau \gg 1$$

$$I_{[a,b]}(t) = \begin{cases} 1 & a \leq t \leq b \\ 0 & \text{otherwise} \end{cases}$$

I is the indicator function

Haar

$$(\Delta v)_{Haar} = \frac{2}{\Delta t} \left[\int_t^{t+\Delta t/2} v(t') dt' - \int_{t-\Delta t/2}^t v(t') dt' \right]$$

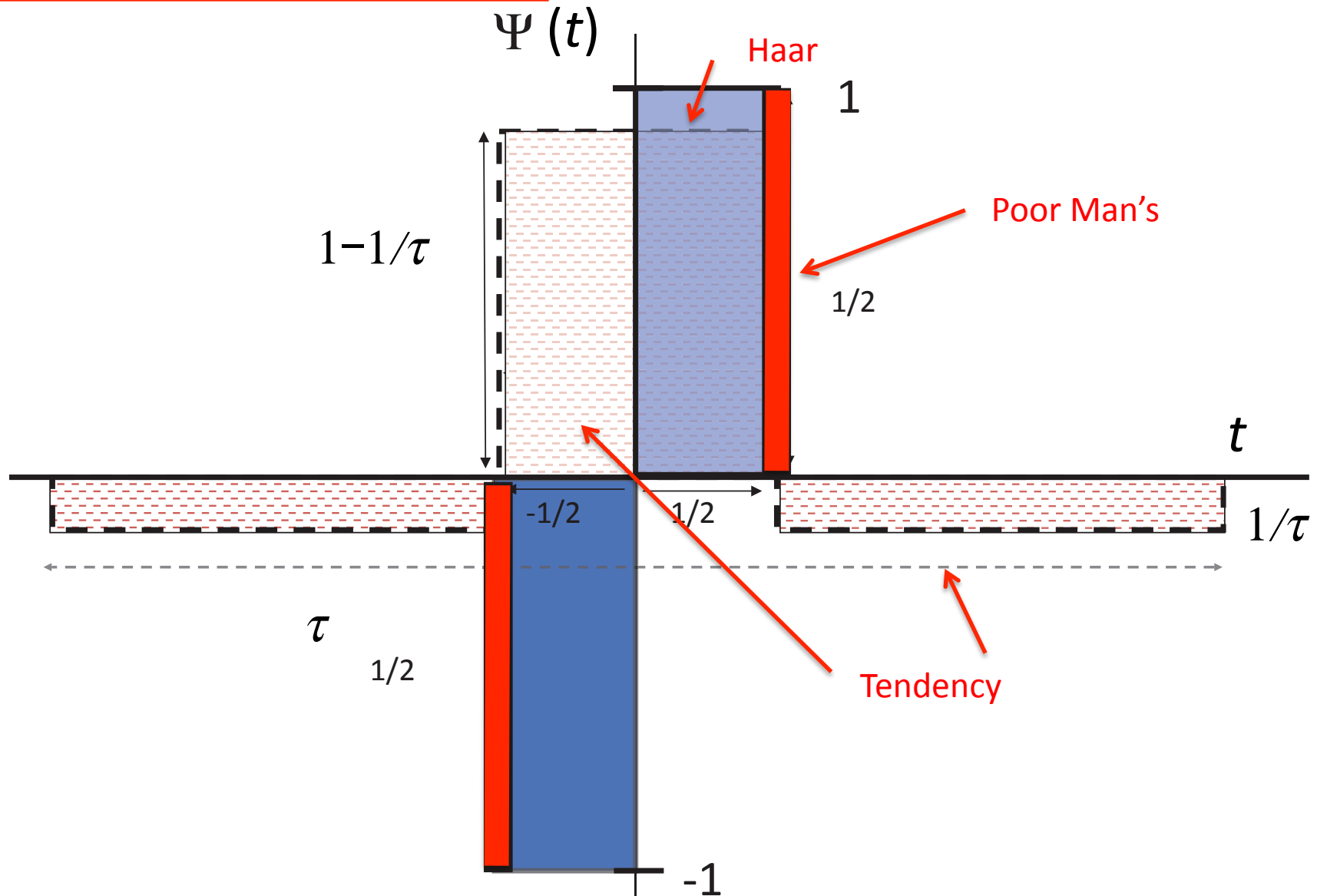
$$\Psi(t) = \begin{cases} 1/2; & 0 \leq t < 1/2 \\ -1/2; & -1/2 \leq t < 0 \\ 0; & \text{otherwise} \end{cases}$$

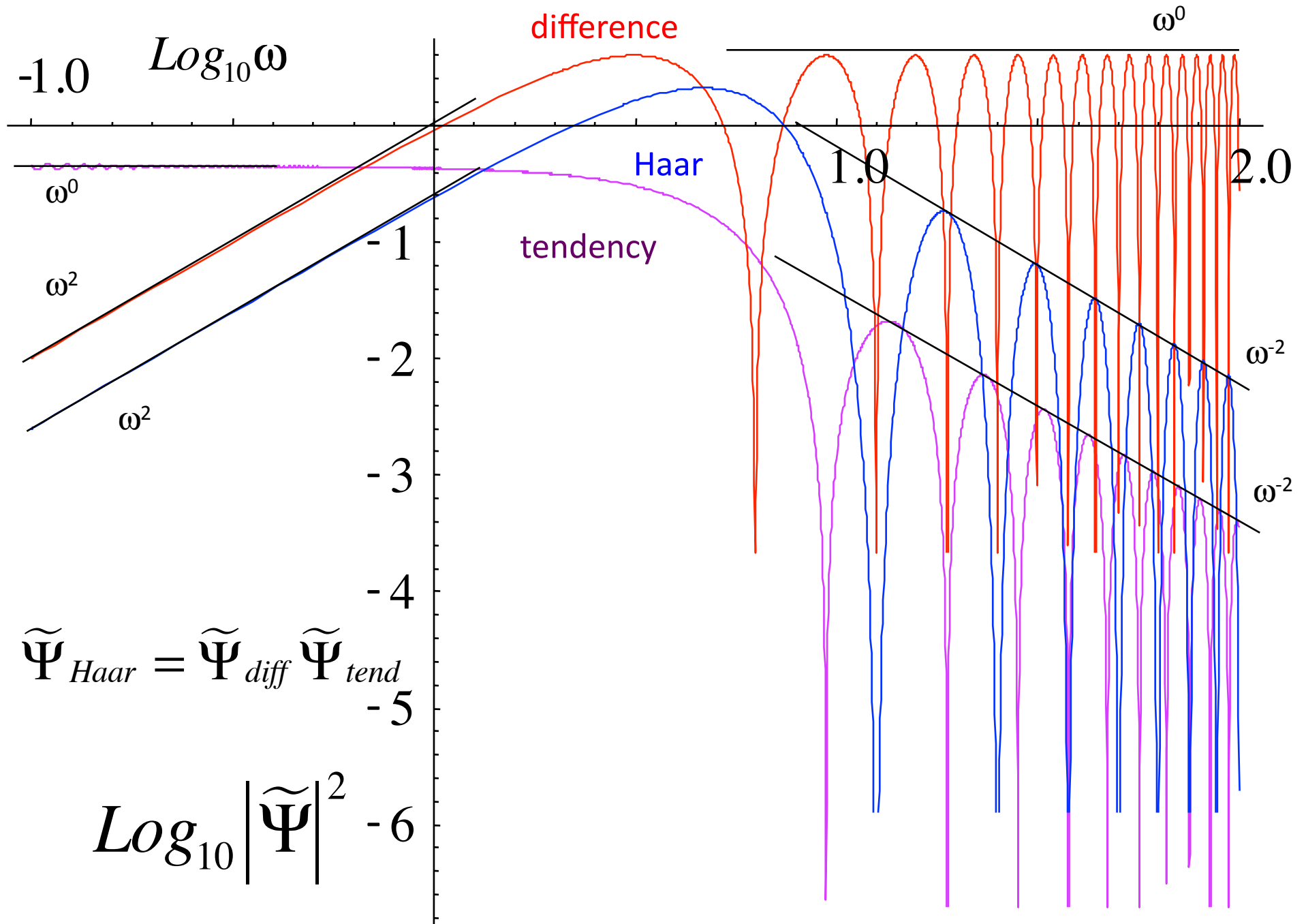
Relation between them:

$$(\Delta v)_{Haar} = (\Delta(\Delta v)_{tend})_{diff}$$

Haar, tendency and poor man's wavelets

$$\Delta v(\Delta t) = \frac{1}{\Delta t} \int v(t') \Psi\left(\frac{t'-t}{\Delta t}\right) dt'$$





Convergence of fluctuation variance

Fluctuations:
$$\Delta v(\Delta t) = \frac{1}{\Delta t} \int v(t') \Psi\left(\frac{t'-t}{\Delta t}\right) dt'$$

Fourier transforms
$$\widetilde{\Delta v}(\omega\Delta t) = \widetilde{v}(\omega) \widetilde{\Psi}(\omega)$$

Ensemble averaging of modulus squared:
$$\langle |\widetilde{\Delta v}(\omega\Delta t)|^2 \rangle = \langle |\widetilde{v}(\omega)|^2 \rangle |\widetilde{\Psi}(\omega)|^2$$

Spectra
$$E_{\Delta v}(\omega) = E_v(\omega) |\widetilde{\Psi}(\omega)|^2 \quad (\Delta t=1)$$

For scaling processes
$$E_v(\omega) \approx \omega^{-(1+2H')}$$

$|\widetilde{\Psi}(\omega)|^2 \approx \begin{cases} \omega^{2H_{low}}; & \omega \rightarrow 0 \\ \omega^{2H_{high}}; & \omega \rightarrow \infty \end{cases}$

$$\langle \Delta v^2 \rangle = \int_{-\infty}^{\infty} E_{\Delta v}(\omega) d\omega$$

Converges only if:

$$H_{low} > H' > H_{high}$$

Parseval's theorem

Various wavelets

Name	Wavelet	Frequency domain	low ω	high ω	H' range
Poor man's (first difference)	$\delta(t-1/2) - \delta(t+1/2)$	$2 \sin(\omega/2)$	$\approx \omega$	≈ 0	$0 \leq H' \leq 1$
2 nd difference	$\frac{1}{2}(\delta(t+1/2) + \delta(t-1/2)) - \delta(t)$	$\sin^2(\omega/4)$	$\approx \omega^2$	≈ 0	$0 \leq H' \leq 1$
Tendency	$I_{[-1/2, 1/2]}(t) - \frac{I_{[-\tau/2, \tau/2]}(t)}{\tau}; \quad \tau \gg 1$	$\frac{2}{\omega} \left(\sin\left(\frac{\omega}{2}\right) - \tau^{-1} \sin\left(\frac{\omega\tau}{2}\right) \right)$	$\frac{2 \sin\left(\frac{\omega\tau}{2}\right)}{\omega\tau} \approx 0; \quad \omega\tau \gg 1$	$\approx \omega^{-1}$	$-1 \leq H' \leq 0$
Haar	$\psi(t) = \begin{cases} 1/2; & 0 \leq t < 1/2 \\ -1/2; & -1/2 \leq t < 0 \\ 0; & \text{otherwise} \end{cases}$	$2i\omega^{-1} \sin^2\left(\frac{\omega}{4}\right)$	$\approx \omega$	$\approx \omega^{-1}$	$-1 \leq H' \leq 1$
Quadratic Haar	$\psi(t) = \begin{cases} -1/3 & 1/3 < t < 1 \\ 2/3; & -1/3 \leq t \leq 1/3 \\ -1/3; & -1 \leq t < -1/3 \\ 0; & \text{otherwise} \end{cases}$	$\frac{2}{3\omega} \left(3 \sin\frac{\omega}{3} - \sin\omega \right)$	$\approx \omega^2$	$\approx \omega^{-1}$	$-1 \leq H' \leq 2$
First derivative Gaussian	$\Psi(t) \propto \frac{d}{dt} e^{-t^2/2}$	$\omega e^{-\omega^2/2}$	$\approx \omega$	$e^{-\omega^2/2}$	$-\infty \leq H' \leq 1$
Mexican Hat	$\Psi(t) \propto \frac{d^2}{dt^2} e^{-t^2/2}$	$\omega^2 e^{-\omega^2/2}$	$\approx \omega^2$	$e^{-\omega^2/2}$	$-\infty \leq H' \leq 2$

Range of exponents over which average fluctuations at scale Δt corresponds to frequency $1/\Delta t$

Fluctuation $\langle \Delta I \rangle = \langle \varphi \rangle \Delta t^H = \text{constant}$

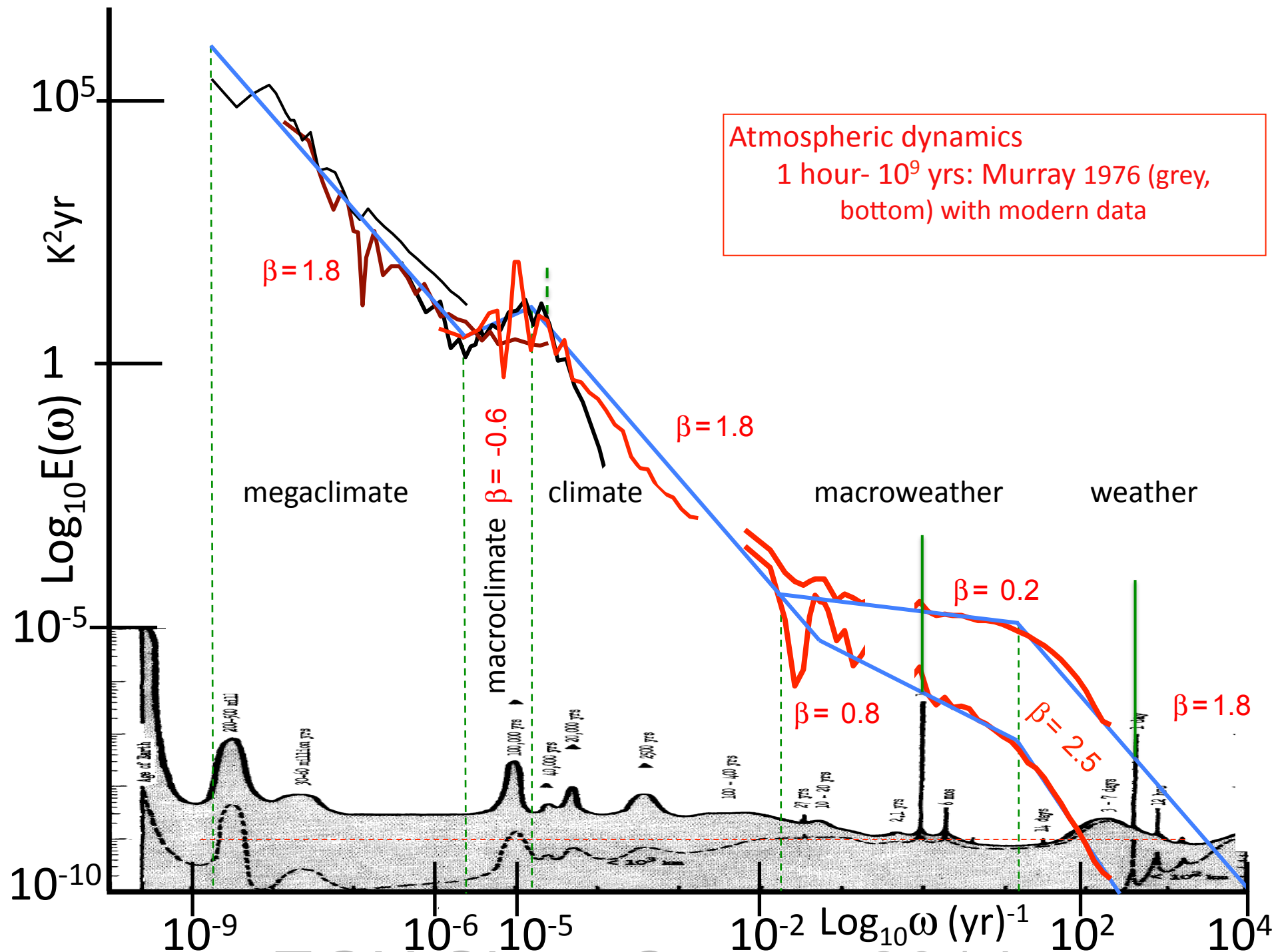
$E(\omega) = \langle |\tilde{I}(\omega)|^2 \rangle = \omega^{-\beta}$

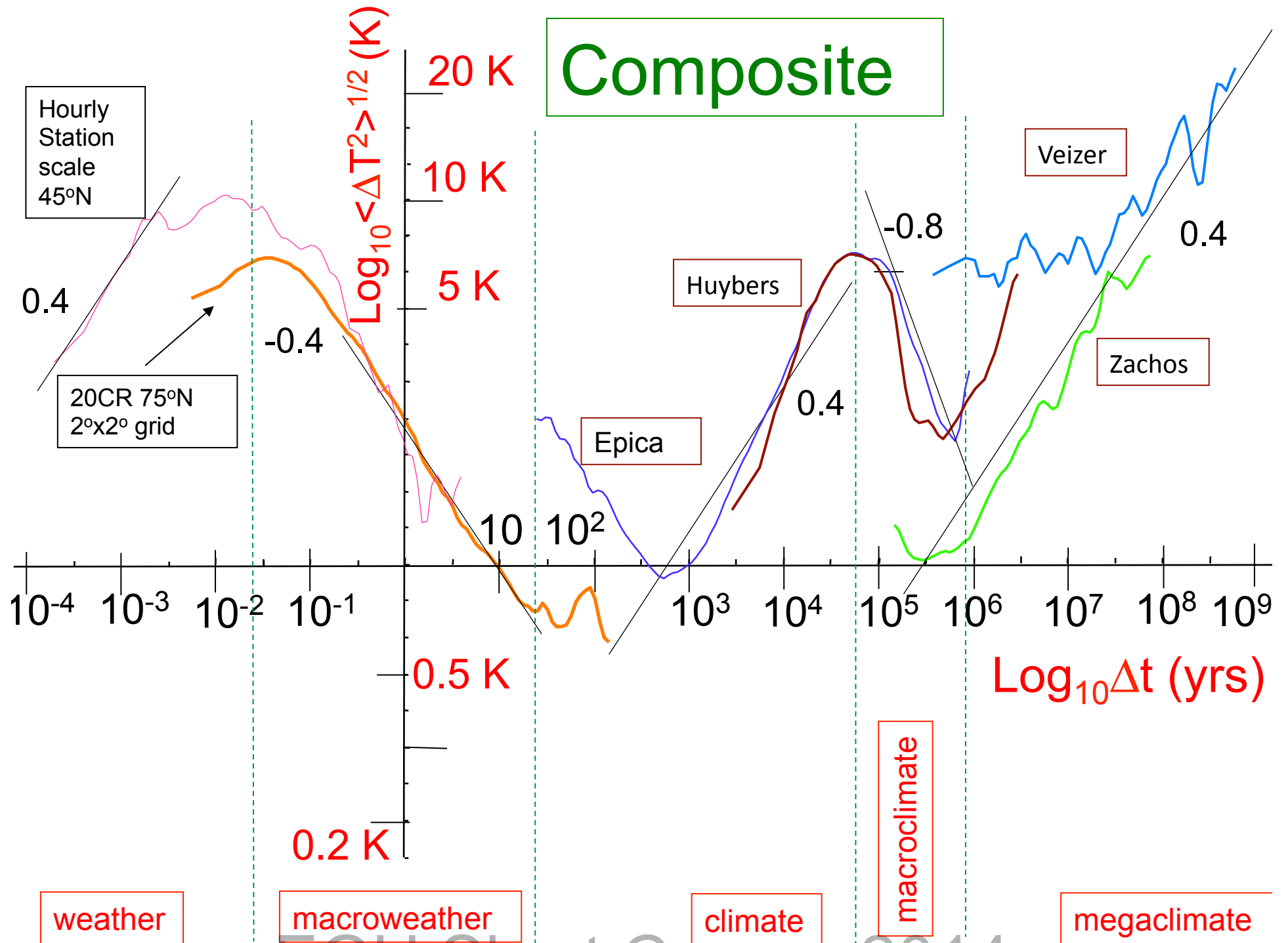
$\beta = 1 + 2H - K(2)$

Multifractal "correction"
 $H' = H - K(2)/2$

Simple interpretation

Statistic	Range of H	Range of β	Comment
Spectrum	$-\infty < H < \infty$	$-\infty < \beta < \infty$	$E(\omega) \approx \omega^{-\beta}$
Difference	$0 < H < 1$	$1 < \beta + K(2) < 3$	"Poor man's wavelet"
Tendency Fluctuation	$-1 < H < 0$	$-1 < \beta + K(2) < 1$	Average with overall mean removed (standard deviation= "Climactogram", also called the "Aggregated Standard Deviation")
Haar	$-1 < H < 1$	$-1 < \beta + K(2) < 3$	Difference of means of first and second halves of interval
Detrended Fluctuation Analysis (DFA, polynomial order n)	$-1 < H < (n+1)$	$-1 < \beta + K(2) < 3+2n$	Also multifractal extension (MFDFA), usually linear: $n=1$, Not a wavelet
Mexican Hat Wavelet	$-\infty < H < 2$	$-\infty < \beta + K(2) < 5$	2 nd Derivative of a Gaussian
Generalized Haar	$-m < H < n$	$1-2m < \beta + K(2) < 3+2n$	Interpretation not simple





$$\langle \Delta T(\Delta t) \rangle \propto \Delta t^H$$

$$H \approx 0.4$$

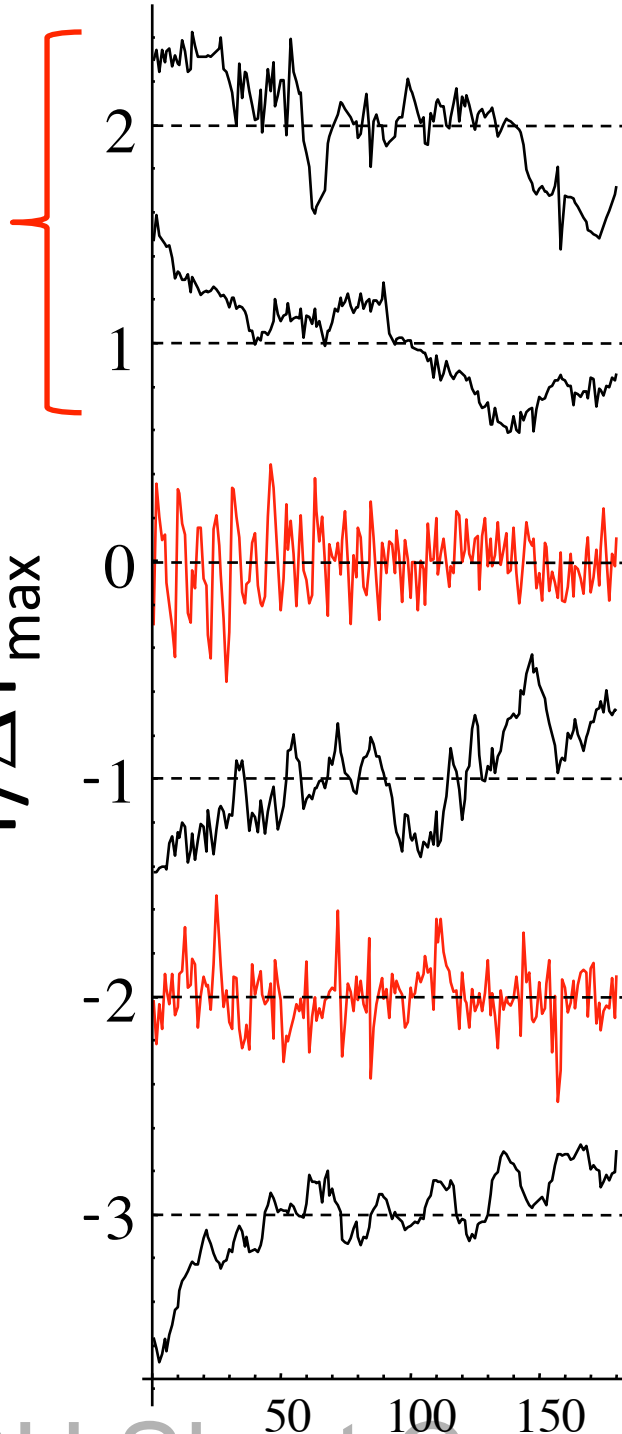
$$H \approx -0.8$$

$$H \approx 0.4$$

$$H \approx -0.4$$

$$H \approx 0.4$$

$T/\Delta T_{\max}$



Megaclimate

Veizer: 290 Mys - 511Myrs BP (1.23Myr)

Megaclimate

Zachos: 0-67 Myrs (370 kyr)

Macroclimate

Huybers: 0-2.56 Myrs (14 kyrs)

Climate

Epica: 25-97 BP kyrs (400 yrs)

Macroweather

Berkeley: 1880-1895 AD (1 month)

Weather

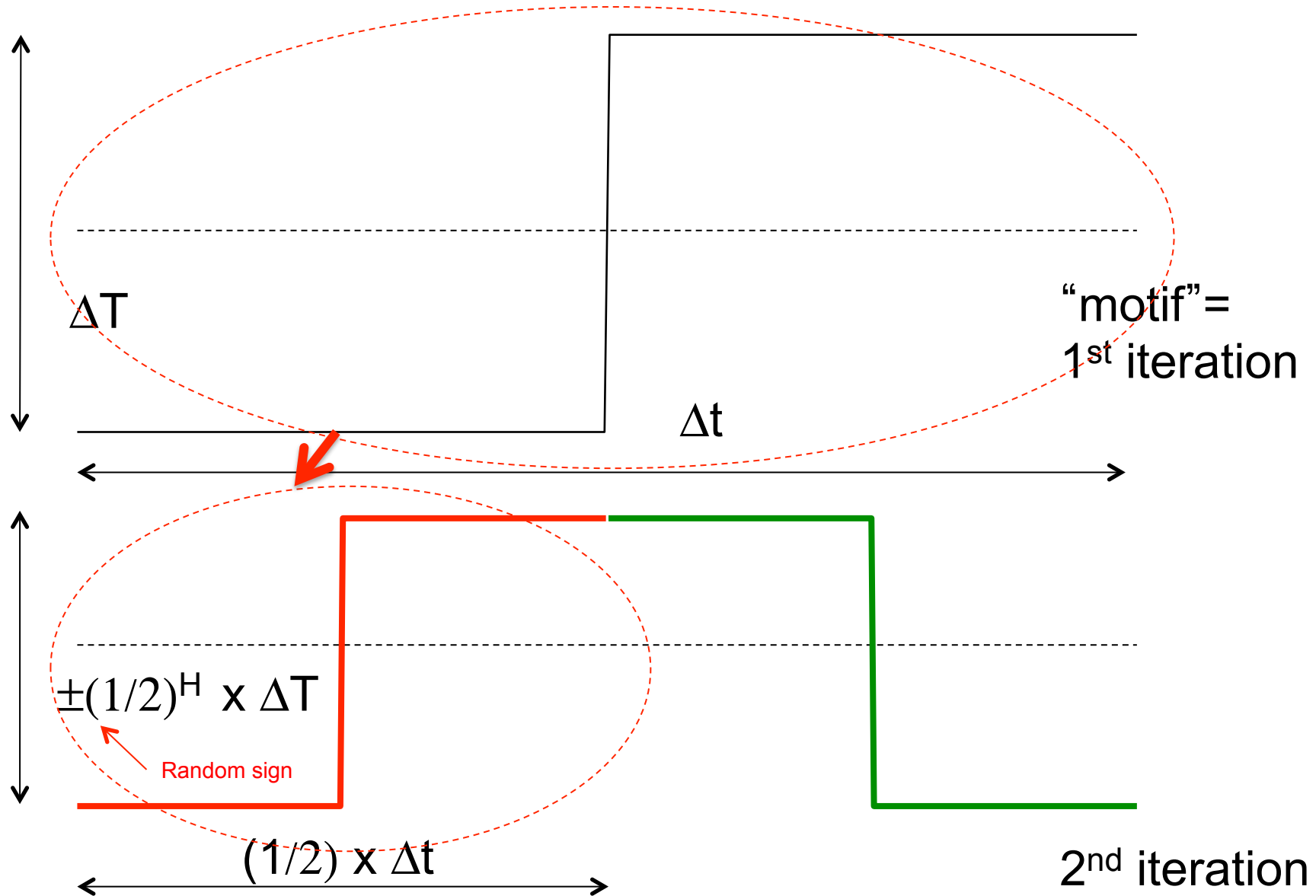
Lander Wy.: July 4-July 11, 2005 (1 hour)

Understanding the Fluctuation exponent H

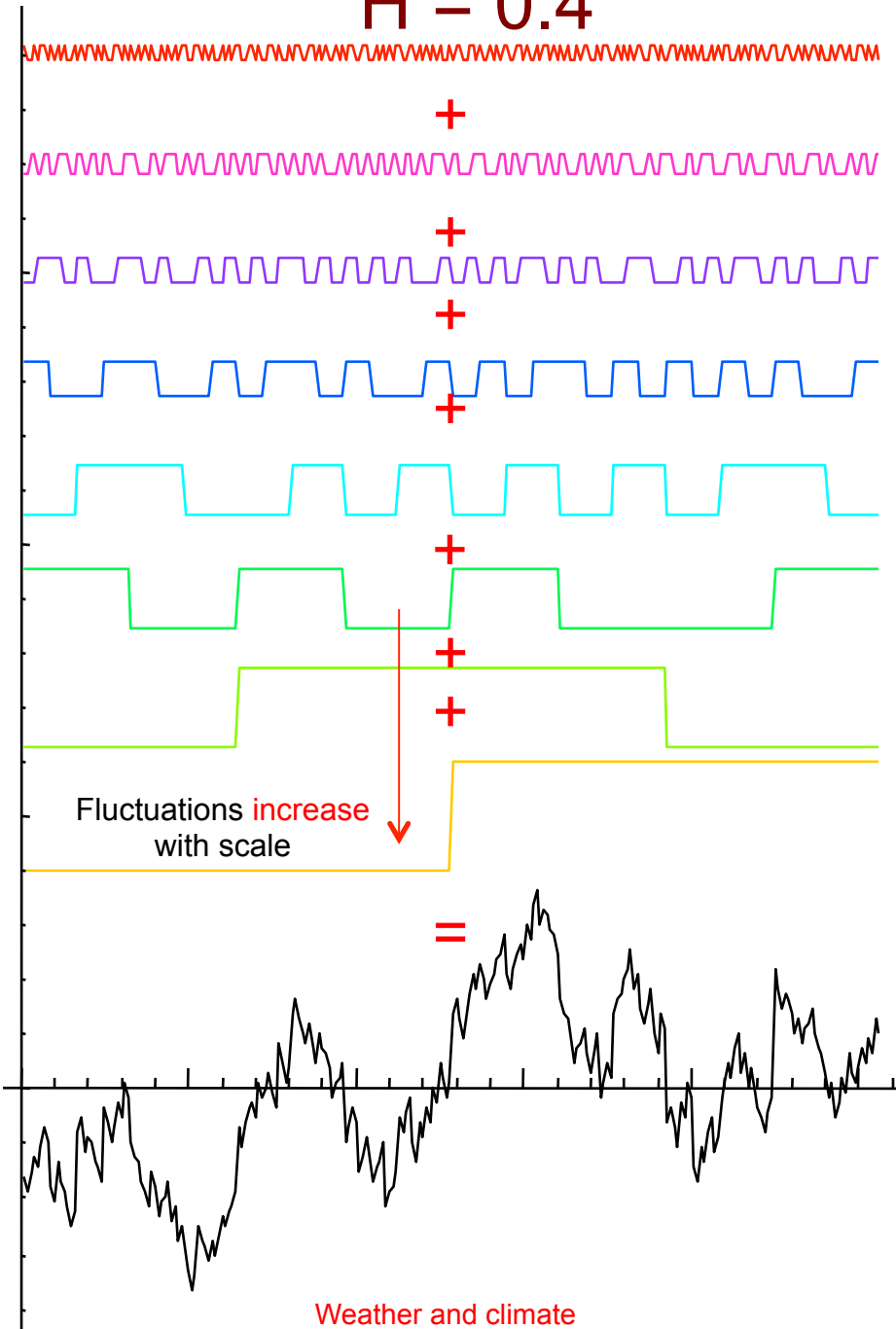
THE HALLAM MODEL

(Lovejoy 2013)

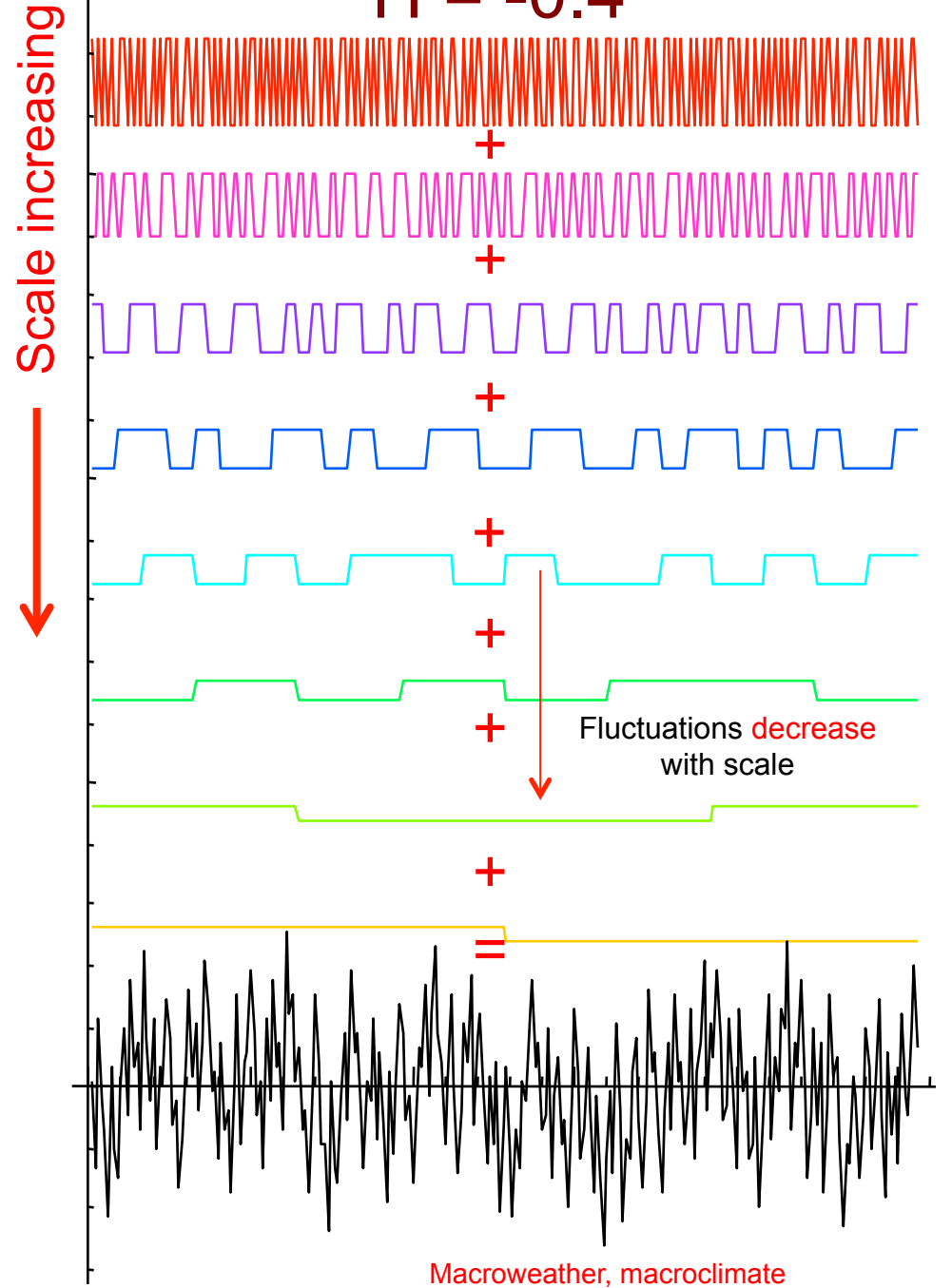
(fractal dimension = $2-H$)



$H = 0.4$



$H = -0.4$



End Part 1